

**PSG COLLEGE OF ARTS & SCIENCE  
(AUTONOMOUS)**

**BSc DEGREE EXAMINATION MAY 2025  
(Fifth Semester)**

Branch – **MATHEMATICS WITH COMPUTER APPLICATIONS**

**MAJOR ELECTIVE COURSE- I: ASTRONOMY**

Time: Three Hours

Maximum: 50 Marks

**SECTION-A (5 Marks)**

Answer **ALL** questions

**ALL questions carry EQUAL marks**

(5 x 1 = 5)

- 1 The period of one complete rotation of the earth about its axis or the period of one apparent rotation of the celestial sphere about the celestial axis is called a .....  
 (i) sidereal noon (ii) sidereal night  
 (iii) astronomical clock (iv) sidereal day
- 2 The terrestrial sphere is divided into a number of regions called the zones of earth by small circle parallel to the equator. These circles are called ..... circles.  
 (i) longitude (ii) latitude  
 (iii) centric (iv) arctic
- 3 Which of the following is the relation between horizontal parallax and angular radius of a body?  
 (i)  $\frac{P}{\theta} = \frac{a}{r}$  (ii)  $\frac{P}{\theta} = \frac{r}{a}$  (iii)  $\frac{P}{\theta} = \frac{1}{r}$  (iv)  $\frac{1}{\theta} = \frac{a}{r}$
- 4 The direction in which a body is seen from the centre of the sun is called its .....  
 (i) heliocentric longitude (ii) geocentric direction  
 (iii) heliocentric direction (iv) mean anomaly
- 5 What is the epact of the year 1952?  
 (i) 2 (ii) 3 (iii) 5 (iv) 4

**SECTION - B (15 Marks)**

Answer **ALL** Questions

**ALL Questions Carry EQUAL Marks**

(5 x 3 = 15)

- 6 a If  $a$  is sun's altitude in the prime vertical at a place of latitude  $\phi$  and  $\lambda$  is its longitude, prove that  $\sin \phi = \sin \lambda \sin \omega \operatorname{cosec} \alpha$   
 OR  
 b Prove that at any instant the sidereal time is given by the right ascension plus west hour angle of any celestial body expressed in time units at that instant.
- 7 a Find analytically the conditions for perpetual day and night.  
 OR  
 b Find an expression for Dip.
- 8 a The altitude of a star is observed and found to be the angle whose sine is  $5/13$ . Calculate the true position of the star, assuming the amount of refraction at an altitude of  $45^\circ$  to be  $58''2$ .  
 OR  
 b Determine the horizontal parallax of moon by meridian observations.

**Cont...**

- 9 a Derive Kepler's third law from Newton's law of gravitation.  
OR  
b Derive Kepler's equation.
- 10 a At the place on the Tropic of cancer, the moon's phase at the time of sun set on 21<sup>st</sup> March was 1/4. show that the altitude of the moon was then 60° assuming the moon to move on the ecliptic.  
OR  
b Find the minimum number of eclipses in a year.

**SECTION -C (30 Marks)**

Answer ALL questions

ALL questions carry EQUAL Marks

(5 x 6 = 30)

- 11 a Trace the changes in the co-ordinates of the sun in the course of a year.  
OR  
b i) At a place in north latitude  $\phi$  two stars A and B (declinations  $\delta$  and  $\delta_1$  respectively) rise at the same moment and A transits when B is setting, prove that  $\tan \phi \tan \delta = 1 - 2 \tan^2 \phi \tan^2 \delta_1$ .  
ii) In a place of latitude 45° shows that the interval between a star's rising and the instant when it is due west is constant.
- 12 a If the declination of the sun changes from  $\delta^\circ$  to  $\delta^\circ + s''$  between sunrise and sunset, show that the afternoon will be longer than the forenoon by  $\frac{\sec \delta \sin \phi}{\sqrt{\cos^2 \phi - \sin^2 \delta}} \frac{s}{15}$  seconds,  $\phi$  being the latitude of the place.  
OR  
b Find the duration of twilight.
- 13 a i) State the laws of refraction  
ii) If  $r''$  be the horizontal refraction, show that the point on the horizon where the sun rises is shifted by  $r'' \sin \phi \sqrt{\sec(\phi - \delta) + \sec(\phi + \delta)}$  where  $\delta$  is the declination of the sun and  $\phi$ , the latitude of the place.  
OR  
b Discuss the changes in R.A. and declination of a body due to geocentric parallax.
- 14 a Obtain the Newton's deductions from Kepler's laws.  
OR  
b If  $t_1, t_2$  are the hourly variations in the equation of time when the sun is at perigee and apogee, show that the eccentricity of earth's orbit is nearly  $\frac{t_1 - t_2}{t_1 + t_2} \tan^2(\omega/2)$ , assuming that the equinoctial line to be perpendicular to the apse line of earth's orbit.
- 15 a Discuss various kinds of lunar librations.  
OR  
b Determine the maximum and minimum number of eclipses possible near a node of the lunar orbit.