

**Statistical Mechanics**  
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**Lecture - 65**  
**Ultra Relativistic Bose Gas Stefan Boltzmann Law**

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Ultra relativistic Bose Gas

$$\epsilon = cp = \underline{hck}$$

$$dk = \frac{d\epsilon}{hc}$$

$$\ln Q_{\eta} = -\eta \sum_k \ln(1 - \eta z e^{-\beta \epsilon})$$

$$\ln Q_{+1} = -\sum_k \ln(1 - z e^{-\beta \epsilon})$$

$$Z \rightarrow \int \frac{V d^3k}{(2\pi)^3} = \frac{4\pi V}{(2\pi)^3} \int_0^{\infty} k^2 dk = \frac{4\pi V}{(hc)^3} \int_0^{\infty} \frac{d\epsilon}{(hc)} \frac{\epsilon^2}{(hc)^2}$$

$$= \frac{4\pi g V}{(hc)^3} \int d\epsilon \epsilon^2$$

$$= \frac{4\pi g V}{(hc)^3} \int d\epsilon \epsilon^2$$

$$g(\epsilon) = \frac{4\pi g V}{(hc)^3} \epsilon^2$$



So, now, that we have looked at an ideal Bose gas confined Bose gas, we will take up the final example in this system. So, what we want to look at is the Ultra Relativistic Bose Gas where my epsilon the energy is given by C times p, which is h bar C k. And I start with the partition function ln Q eta; the general expression was minus eta sum over k ln 1 minus eta z e to the power minus beta epsilon.

Now, here eta is plus 1, eta is equal to plus 1 for a bosons. So, I have Q plus is going to be minus sum over k ln 1 minus Z e to the power minus beta epsilon. The sum over k as usual is converted to this density of state, as V over twice by whole cube integral k square over dk in

three dimension and since I have this expression, I can write down  $dk$ , as  $d\epsilon$  over  $h c$ .

So, that I have  $V$  over there has to be a  $gV$  here and a  $4\pi$  over here. We will substitute the value of  $g$  a little later, let us keep it  $4\pi$  times  $gV$  divided by  $2\pi$  whole cube and then I have integration of  $d\epsilon$  over  $h c$ . And I have  $h c$  Square in terms of replacing  $k$  sorry, this has to be  $\epsilon^2$  over  $h c$  whole square.

So, that I have  $4\pi gV$  divided by  $2\pi h c$  whole cube integration  $d\epsilon \epsilon^2$ , which gives me  $4\pi gV$  divided by  $h c$  whole cube integration  $d\epsilon \epsilon^2$ . So, that the density of state, we will straight forward write down as  $4\pi gV h c$  whole cube  $\epsilon^2$ .

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$$g(\epsilon) = \frac{4\pi g V}{(hc)^3} \epsilon^2$$

$$\ln Q_+ = - \int d\epsilon g(\epsilon) \ln(1 - z e^{-\beta\epsilon}) = - \frac{4\pi g V}{(hc)^3} \int_0^\infty d\epsilon \epsilon^2 \ln(1 - z e^{-\beta\epsilon})$$

$$= \frac{4\pi g V}{(hc)^3} \left[ \int_0^\infty d\epsilon \epsilon^2 \ln(z e^{-\beta\epsilon}) - \int_0^\infty d\epsilon \epsilon^2 \frac{z e^{-\beta\epsilon}}{1 - z e^{-\beta\epsilon}} (\beta\epsilon) \right]$$

$$\ln Q_+ = \frac{4\pi g V}{(hc)^3} \beta \int_0^\infty d\epsilon \frac{\epsilon^3}{z^{-1} e^{\beta\epsilon} - 1}$$

$$\beta P V = \ln Q_+ = \frac{4\pi g V}{(hc)^3} \beta \int_0^\infty d\epsilon \frac{\epsilon^3}{z^{-1} e^{\beta\epsilon} - 1}$$



Now, this means, that  $\ln$  of  $Q$  plus is going to be minus integral, let us write down the factors ok. So,  $d\epsilon$ ,  $g\epsilon$   $1 - Z e$  to the power minus  $\beta\epsilon$   $\ln$  of this, which is going to be minus  $4\pi gV$  over  $hC$  whole cube  $d\epsilon$   $\epsilon^2 \ln 1 - Z e$  to the power minus  $\beta\epsilon$ .

We carried out these integrals, the way to go ahead is to use integration by parts and we will do it over here also, this is  $\epsilon^3$  by  $3 \ln$  of  $1 - Z e$  to the power minus  $\beta\epsilon$ ,  $0$  to infinity this is  $0$  to infinity minus; I am going to have integration  $d\epsilon$   $\epsilon^3$  by  $3$  derivative of this  $\ln$  is going to give me  $Z e$  to the power minus  $\beta\epsilon$  minus  $Z e$  to the power minus  $\beta\epsilon$ .

And then  $d\epsilon$  of the minus  $\beta\epsilon$ , which is going to give me a minus  $\beta$  factor. So, that this minus and this minus is going to make it a plus. And this minus and this minus is going to make it a plus. And I know that this first term is going to vanish, I am going to have  $4\pi gV$   $hC$  whole cube  $\beta$  no; yeah,  $\beta$  over  $3$  integral  $0$  to infinity  $d\epsilon$   $\epsilon^3 Z$  inverse  $e$  to the power  $\beta\epsilon$  minus  $1$ .

This is  $\ln$  of  $Q$  plus and I know that  $\beta PV$  is going to be  $\ln$  of  $Q$  plus, which is going to be  $4\pi gV$   $hC$  whole cube  $\beta$  over  $3$ ,  $0$  to infinity  $d\epsilon$   $\epsilon^3 Z$  inverse  $e$  to the power  $\beta\epsilon$  minus  $1$ .

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$$\begin{aligned}
 \beta P V &= \ln Q_1 = \frac{4\pi g V}{(hc)^3} \frac{1}{3} \int_0^\infty d\epsilon \frac{\epsilon^3}{z^{-1} e^{\beta\epsilon} - 1} \quad \epsilon=0 \\
 U &= \sum_k \epsilon_k \langle n_k \rangle \\
 U &= \int g(\epsilon) d\epsilon \frac{\epsilon}{z^{-1} e^{\beta\epsilon} - 1} = \frac{4\pi g V}{(hc)^3} \int_0^\infty d\epsilon \frac{\epsilon^3}{z^{-1} e^{\beta\epsilon} - 1} \\
 \mu=0 &\Rightarrow z=1 \\
 P &= \frac{4\pi g}{(hc)^3} \frac{1}{3} \int_0^\infty d\epsilon \frac{\epsilon^3}{e^{\beta\epsilon} - 1} \\
 &= \frac{d(\beta\epsilon)}{\beta} \frac{1}{\beta^3} \frac{(\beta\epsilon)^3}{e^{\beta\epsilon} - 1}
 \end{aligned}$$



The total energy is going to be  $g \epsilon^3 d\epsilon$  and  $\epsilon$  divided by  $Z^{-1} e^{\beta\epsilon} - 1$ . We have seen it several times, because this follows from the relation the total energy is sum over  $k$   $\epsilon_k \langle n_k \rangle$ . This gives me  $4\pi g V / (hc)^3$ . Integration  $d\epsilon$ ;  $g \epsilon^3$  behaves as  $\epsilon^4$  and I have an additional  $\epsilon$  factor here so, this gives me  $\epsilon^5$   $Z^{-1} e^{\beta\epsilon} - 1$ .

Now, if you turn your attention to photons, it is something is very very special. So, what you see over here, that for low excitations like; photons the chemical potential  $\mu$  is equal to 0, because it does not cost you energy to create particles at  $\epsilon=0$  right. So, once  $\mu$  is equal to 0 this implies that  $Z$  is equal to 1.

So, essentially, where the thermodynamic pressure becomes  $4\pi g$ ; the volume factor cancels over here, the  $\beta$  factor cancels over here,  $(hc)^3$  one-third 0 to infinity  $d\epsilon$   $\epsilon^3$

cube  $\epsilon$  to the power  $\beta \epsilon^3$  minus 1. And I can substitute for this 0 to infinity  $d$  of  $\beta \epsilon^3$  divided by  $\beta$   $\beta^3$   $\epsilon^3$  whole cube 1 over  $\beta$  cube  $\epsilon$  to the power  $\beta \epsilon^3$  minus 1.

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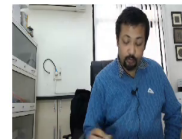
$$P = \frac{4\pi g}{(hc)^3} \frac{1}{3} \int_0^\infty d\epsilon \frac{\epsilon^3}{e^{\beta \epsilon^3} - 1}$$

$$P = \frac{4\pi g}{(hc)^3} \frac{(k_B T)^4}{3} \frac{\Gamma(4)}{\Gamma(4)} \int_0^\infty dx \frac{x^3}{e^x - 1} = \frac{\zeta(4) \Gamma(4)}{\Gamma(4)} = \zeta(4) \Gamma(4)$$

$$P = \frac{4\pi g}{(hc)^3} (k_B T)^4 \zeta(4) \Gamma(4)$$

$$U = \frac{4\pi g V}{(hc)^3}$$

$\Gamma(m) = (m-1)!$   
 $\int_0^\infty \frac{x^{m-1}}{e^x - 1} dx = \frac{1}{(m-1)!} \int_0^\infty \frac{dx x^{m-1}}{e^x - 1}$   
 $\zeta(4) \Gamma(4) = \frac{\pi^4}{15}$



So, that the thermodynamic pressure I can write down;  $4\pi g$  over  $hC$  whole cube  $k_B T$  raised to the power 4 by 3, 0 to infinity  $dx$   $x$  cube over  $e$  to the power  $x$  minus 1. Now, the value of this integral I know. The value of this integral is zeta 4 times gamma 3 am I right, sorry this has to be gamma 4.

Because I know, from our earlier discussion that  $F_m$  plus of  $Z$  is 1 over  $M$  minus 1 factorial  $dx$   $x$  to the power  $m$  minus 1  $z$  inverse  $e$  to the power  $x$  minus 1. And gamma of  $m$  is going to be  $m$  minus 1 factorial. So, essentially I multiply this by gamma  $m$  sorry, gamma 4 and divided by gamma 4 and I realize, that this quantity that I see over here is equivalent to  $F_4$  plus of 1, which is zeta times 4.

So, that the thermodynamic pressure is  $\frac{4\pi g}{(hc)^3} (k_B T)^4$  whole cube  $k_B T$  raised to the power 4 gamma 4 zeta 4. You should know, that gamma 4 times zeta 4 is  $\frac{\pi^4}{15}$ . We will later on use this one. Now, the internal energy, it follows is  $\frac{4\pi g V}{(hc)^3} (k_B T)^4$  whole cube. Once again if you look at the structure of this internal energy I am going to have 0 to infinity.

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$$P = \frac{4\pi g}{(hc)^3} \frac{(k_B T)^4}{3} \int_0^\infty dx \frac{x^3}{e^x - 1} = \frac{4\pi g}{(hc)^3} \frac{(k_B T)^4}{3} \zeta(4) \Gamma(4)$$

$$\zeta(4) \Gamma(4) = \frac{\pi^4}{15}$$

$$P = \frac{4\pi g}{(hc)^3} \frac{(k_B T)^4}{3} \zeta(4) \Gamma(4)$$

$$U = \frac{4\pi g V}{(hc)^3} \int_0^\infty \frac{d^3 p}{p^3} \frac{(p c)^3}{e^{p c / k_B T} - 1} = \frac{4\pi g V}{(hc)^3} (k_B T)^4 \zeta(4) \Gamma(4)$$

$$\frac{1}{3} \frac{U}{V} = P \quad \text{Ultra relativistic Gas.}$$

$$P \sim T^4 \quad \frac{U}{V} \sim T^4$$



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$$\beta P V = \ln Q_T = \frac{4\pi g V}{(hc)^3} \int_0^\infty d\epsilon \frac{\epsilon^3}{z^{-1} e^{\beta\epsilon} - 1} \quad \epsilon=0$$

$$U = \int_0^\infty g(\epsilon) d\epsilon = \frac{4\pi g V}{(hc)^3} \int_0^\infty d\epsilon \frac{\epsilon^3}{z^{-1} e^{\beta\epsilon} - 1}$$



$$\mu=0 \Rightarrow z=1$$

$$P = \frac{4\pi g}{(hc)^3} \frac{1}{3} \int_0^\infty d\epsilon \frac{\epsilon^3}{e^{\beta\epsilon} - 1}$$

$$P = \frac{4\pi g}{(hc)^3} \frac{(k_B T)^4}{3} \frac{\Gamma(4)}{\Gamma(4)} \int_0^\infty dx \frac{x^3}{e^x - 1} = \frac{4\pi g}{(hc)^3} \frac{(k_B T)^4}{3} \Gamma(4) = 5/4$$

$$\int_0^\infty \frac{x^m}{e^x - 1} dx = \frac{1}{(m-1)!} \int_0^\infty \frac{x^{m-1}}{e^x - 1} dx$$

$$\Gamma(m) = (m-1)!$$

So, I have always missed out, this integral limits. So, please note that, they always run from 0 to infinity;  $d$  of beta epsilon the standard trick is apply over here, 1 over beta cube epsilon to the power beta h epsilon minus 1 and that is going to be 4 pi gV over hC whole cube k B T raised to the power 4 again it is going to give you gamma 4 zeta 4.

So, if you compare this expression over here and the last expression over here, you see that; U by V I have missed out a factor 3 over here. One-third of U by V is the pressure, which is the case for an ultra relativistic gas, which we have done several times in this course. Additionally, you must note that the pressure goes as T to the power 4 and U goes as T to the power 4 in this case. Now, that I know the energy density even I can write down U by V, the energy density goes as T to the power 4.

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Black body Radiation


$P \sim T^4$      $\frac{U}{V} \sim T^4$      $\frac{1}{3} \frac{U}{V} = 1$      $g(K) \rightarrow g(\epsilon)$

$E = k\omega$      $g(\omega)$

$g(\epsilon) = \frac{4\pi g V}{(hc)^3} \epsilon^2$      $E = k\omega$   
 $\frac{d\epsilon}{d\omega} = k$

$\sum_K \equiv \frac{4\pi g V}{(hc)^3} \int_{k\omega}^2 \frac{d\epsilon}{k\omega^2 k d\omega}$

$\equiv \frac{4\pi g V}{(hc)^3} k^3 \int \omega^2 d\omega$




Now, suppose I have a container, which has a tiny tiny hole and this is filled with photons. So, that these photons come out from this hole. So, essentially what I am trying to look at is; what is called a black body radiation right. For this, I want to write down epsilon as  $h \bar{\omega}$ .

So, I started off with K, I went to g epsilon. So, g of K the density of states and from this, I want to go to g of omega. So, g of epsilon was  $4 \pi g V$  divided by  $hC$  whole cube and then, I had  $\epsilon^2$  square. So, that this sum over K, I had it was equivalent to  $4 \pi g V$  over  $hC$  whole cube integration  $\epsilon^2 d\epsilon$ , do not forget that.



Now,  $\epsilon_{psa}$  is  $\hbar \omega$  and therefore,  $d \epsilon_{psa} d \omega$  is going to be  $\hbar$ . So, this becomes  $\hbar d \omega$  and this becomes  $\hbar^2 \omega^2 d \omega$ . So, that this is;  $4 \pi gV$  over  $h^3 c^3$  whole cube times  $\hbar^2 \omega^2 d \omega$ .

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$$\begin{aligned}
 &= \frac{4\pi g V \hbar^3}{(h c)^3} \int \omega^2 d\omega \\
 &= \frac{4\pi g V \hbar^3}{\hbar^3 c^3 (2\pi)^3} \int \omega^2 d\omega \\
 &= \frac{4\pi g V}{c^3 8\pi^3} = \frac{g V}{2 c^3 \pi^2} \int \omega^2 d\omega \\
 g(\omega) &= \frac{g V}{2 c^3 \pi^2} \omega^2 \\
 g(\omega) &= \frac{V \omega^2}{c^3 \pi^2}
 \end{aligned}$$



Which is equivalent to  $4 \pi gV$  divided by; I have  $h^3$ ,  $h^3 c^3$  and  $2 \pi$  whole cube. This is the part, which follows from  $\hbar$ . So, that and then I have  $\omega^2 d \omega$ . This becomes  $4 \pi gV$ , the  $h^3 h^3$  cancels out,  $c^3 8 \pi^3$ ; this is  $2$  and I have  $\pi^2$  square. So, I have  $gV$  over  $2 c^3 \pi^2$  integration  $\omega^2 d \omega$ .

So, that the density of state in the frequency spectrum is equivalent to  $gV$  over  $2 c^3 \pi^2$  square,  $\omega^2$ . Now, for a photon the degeneracy factor is  $2$ . So, that I can write down

this as;  $v \omega^2$  over  $C^3 \pi^2$  a very nice expression. For this density of state right.

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$2c^3 n^2$

$$g(\omega) = \frac{V \omega^2}{c^3 \pi^2}$$

The number of oscillators within a range of frequency between,  $\omega$  &  $\omega + d\omega$

$$dN(\omega) = \langle n_\omega \rangle g(\omega) d\omega$$

$$= \left( \frac{1}{e^{\beta \hbar \omega} - 1} \right) \frac{V \omega^2}{c^3 \pi^2} d\omega$$

$$\frac{dN(\omega)}{d\omega} = V \frac{\omega^2}{c^3 \pi^2} \frac{1}{e^{\beta \hbar \omega} - 1}$$

$$\frac{1}{V} \frac{dN(\omega)}{d\omega} = \frac{\omega^2}{c^3 \pi^2} \frac{1}{e^{\beta \hbar \omega} - 1}$$



Therefore, it follows that a number of oscillators within a range of frequency between  $\omega$  and  $\omega + d\omega$  is going to be  $dN(\omega)$ , which is going to be average of  $n_\omega$ ,  $g(\omega)$ ,  $d\omega$ . And average of  $n_\omega$ , I know is  $e^{-\beta \hbar \omega}$  minus 1, because  $Z$  is 1;  $g(\omega)$  is  $V \omega^2 / C^3 \pi^2$  times  $d\omega$ .

So, one has  $dN(\omega) d\omega$  as  $v \omega^2$  over  $C^3 \pi^2$   $1 / \beta \hbar \omega$  minus 1. This part the 1 in the bracket essentially follows from this average of  $n_\omega$ , which I know average of  $n_k$  is  $1 / Z$  inverse  $e^{-\beta \hbar \omega}$  minus 1 and  $Z$  is 1 for a photon gas,  $\epsilon$  is  $\hbar \omega$  so, I have this expression right.

Now, I can bring the volume in the denominator. So, that I have 1 by V dN omega d omega is going to be omega square C cube, pi square 1 over e to the power beta h omega minus 1 correct.

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$$\left(\frac{1}{V} \frac{dN(\omega)}{d\omega}\right)' = \frac{\omega}{c^3 \pi^2} \frac{1}{e^{\beta \hbar \omega} - 1}$$

$$\frac{N}{V} = n$$

Density of oscillators which are between  $\omega$  &  $\omega + d\omega$

$$\frac{dN(\omega)}{d\omega} = \frac{\omega^2}{c^3 \pi^2} \frac{1}{(e^{\beta \hbar \omega} - 1)}$$

$$dU(\omega) = \epsilon_\omega \langle n_\omega \rangle g(\omega) d\omega$$

$$= \hbar \omega \frac{1}{e^{\beta \hbar \omega} - 1} \frac{V \omega^2}{c^3 \pi^2} d\omega$$

$$\frac{1}{V} dU(\omega) = \frac{\hbar}{c^3 \pi^2} \frac{\omega^3}{e^{\beta \hbar \omega} - 1}$$



Let us see, then look at the left hand side. N by V is the number density small n. So, therefore, the density of oscillators, which are between so, density of oscillators, which are between omega and omega plus d omega, which has frequencies between omega and omega plus d omega is dn omega d omega is given by this expression omega square C cube pi square 1 over e to the power beta h bar omega. Good.

Now, I want to calculate the amount of energy that is contained within this frequency interval. So, now, I want to know the energy density contained within this frequency interval omega and omega plus d omega. So, one can also write down dU of omega, which is that amount of

energy, which is contained within that is going to be;  $\epsilon_{psa} \omega$ , average of  $n \omega$ ,  $g \omega$  and  $d \omega$ .

$\epsilon_{psa} \omega$  is  $\hbar \omega$  and  $n \omega$  is  $1$  over  $e$  to the power  $\beta \hbar \omega$  minus  $1$  and then, you have  $g \omega$ , which is  $\omega^2 C^3 \pi^2 d \omega$ . There is a  $V$  factor over here  $\omega^2 V$ . So, the energy density; so,  $1$  by  $V dU$  of  $\omega$  is going to be  $\hbar C^3 \pi^2 \omega^3$  divided by  $e$  to the power  $\beta \hbar \omega$  minus  $1$ .

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$$\frac{dn(\omega)}{d\omega} = \frac{\omega}{c^3 n^2} \frac{1}{(e^{\beta \hbar \omega} - 1)}$$

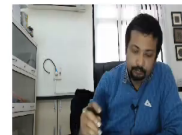
$$dU(\omega) = \epsilon_{\omega} \langle n_{\omega} \rangle g(\omega) d\omega$$

$$= \hbar \omega \frac{1}{e^{\beta \hbar \omega} - 1} \frac{V \omega^2}{c^3 n^2} d\omega$$

$$\frac{1}{V} dU(\omega) = \frac{\hbar}{c^3 n^2} \frac{\omega^3}{e^{\beta \hbar \omega} - 1}$$

$$u = \frac{U}{V}$$

$$\frac{du(\omega)}{d\omega} = \frac{\hbar}{c^3 n^2} \frac{\omega^3}{(e^{\beta \hbar \omega} - 1)}$$



So, clearly this means; that the energy density small  $u$  of  $\omega$ , where small  $u$  is defined as capital  $U$  divided by  $V$  the density, divided by  $d \omega$  is going to be  $\hbar C^3 \pi^2 \omega^3$   $e$  to the power  $\beta \hbar \omega$  minus  $1$ . Now, one has to go back to the original problem and look at this; that I have a box, which is filled with photons and I have a tiny tiny hole here from which the photons emerge.

So, I am going to want to know the flux, and the flux; if these photons move with the velocity  $C$ , the flux is going to be  $n$  over one-fourth  $n$  times  $C$  right where it is understood; that this is going to be  $n$  omega times. So, now this is the energy density. So, if you want to ask, how what is the amount of energy which is contained within a frequency interval  $\omega$  and  $\omega + d\omega$ ; the answer would be  $du$   $d\omega$  times  $d\omega$ .

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$$dU(\omega) = E_{\omega} \langle n_{\omega} \rangle g(\omega) d\omega$$

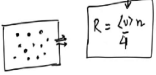
$$= \left( \hbar\omega \right) \frac{1}{e^{\beta\hbar\omega} - 1} \frac{V \omega^2}{c^3 n^2} d\omega$$

$$\frac{1}{V} dU(\omega) = \frac{\hbar}{c^3 n^2} \frac{\omega^3}{e^{\beta\hbar\omega} - 1}$$

$$\frac{du(\omega)}{d\omega} = \frac{\hbar}{c^3 n^2} \frac{\omega^3}{(e^{\beta\hbar\omega} - 1)}$$

$$\text{Energy flux} = \frac{c}{4} \hbar\omega \frac{dn}{d\omega}$$

$\omega, \omega + d\omega$   
 $\frac{du}{d\omega} d\omega$   
 $\frac{dn}{d\omega} d\omega$   
 $R = \frac{c}{4} n$





And similarly, the number of oscillators you have within the frequency interval  $\omega$  and  $\omega + d\omega$ , you have  $dn$   $d\omega$  times  $d\omega$ . Now, the original question that we wanted to ask was that; we have this problem where we have a box and there is a small hole from which photons are coming out, right.

So, this is the radiation black body radiation, which you are looking at. So, I want to know the energy flux correct. So, here this is the classical problem of effusion and the answer to that is

the number density, the number flux is going to be average  $V$  divided by 4 times  $n$ . This is for the classical result.

In our case, we adopt this relation and identify that we are looking at the frequency interval,  $\omega$  and  $\omega + d\omega$ . And therefore, we have number of oscillators, which is  $dn$   $d\omega$  and each of these oscillators carry an energy  $\hbar \omega$ , which we have already identified.

Therefore the energy flux  $du$   $d\omega$ ; so, we will write down this as, the energy flux that is coming out of this small tiny hole is going to be  $\hbar \omega$  sorry, the average velocity is the velocity of light, which is  $C$  by 4 and I have  $dn$   $d\omega$ . This is per frequency interval right.

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$$u = \frac{U}{V}$$

$$\frac{d u(\omega)}{d \omega} = \frac{\hbar}{c^3 n^2} \frac{\omega^3}{(e^{\beta \hbar \omega} - 1)}$$

$\frac{du}{d\omega} \left( \frac{dn}{d\omega} \right) d\omega$

$R = \frac{c \omega n}{4}$

Energy flux =  $\frac{c}{4} \hbar \omega \frac{dn}{d\omega} \rightarrow$  per frequency interval.

$\Delta E = \frac{c}{4} \hbar \omega \frac{dn}{d\omega} d\omega$

$$dE = \frac{c}{4} \hbar \omega \frac{\omega^2}{c^3 n^2} \frac{1}{e^{\beta \hbar \omega} - 1} d\omega$$

$$= \frac{\hbar}{4 n^2 c^2} \frac{\omega^3}{e^{\beta \hbar \omega} - 1} d\omega$$



So, this is per frequency interval. Once again, if you want to know how much energy is coming out within this; so, let us say  $\Delta E$  is going to be  $C^4$ ,  $h$  bar by  $\omega$   $d\omega$   $d\omega$  times  $d\omega$ .

This is the amount of energy, which is coming out within a frequency interval of  $\omega$  and  $\omega + d\omega$  which we write down as  $C^4$ ,  $h$  bar  $\omega$  and then I have  $d\omega$   $d\omega$   $d\omega$ , which we had before which was this expression becomes  $\omega^3 C^4$  and I have  $\pi^2$   $1$  over  $e$  to the power  $\beta h$  bar  $\omega$  minus  $1$  and this gives me  $\omega^3$  over  $ok$ ; so, first we write down as  $h$  bar,  $h$  bar  $1$  factor of  $C$  cancels with  $1$  factor also t give you  $C^4$ .

So, you have  $4\pi^2 C^4$  and then you have  $\omega^3$ , this  $\omega$  and this  $\omega^2$  square here and then you have  $e$  to the power  $\beta h$  bar  $\omega$  minus  $1$ . So, if you were to measure the amount of energy contained, which is coming out of this tiny pore within a frequency interval of  $\omega$  and  $\omega + d\omega$ , then you will see that this is the  $2$  amount  $\Delta E$  times  $d\omega$ .

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$$dE = \frac{c}{4} k \omega \frac{dn}{d\omega} d\omega$$

$$dE = \frac{h}{4} \frac{k \omega^3}{\pi^2 c^2} \frac{1}{e^{\beta h \omega} - 1} d\omega$$

$$dE = \frac{h}{4 \pi^2 c^2} \frac{\omega^3}{e^{\beta h \omega} - 1} d\omega$$

$$E = \frac{h}{4 \pi^2 c^2} \int_0^{\infty} d\omega \frac{\omega^3}{e^{\beta h \omega} - 1}$$

$$= \frac{h}{4 \pi^2 c^2} \int_0^{\infty} \frac{dx}{(\beta h)^3} \frac{1}{e^x - 1} \frac{x^3}{\beta^3}$$

$$= \frac{h}{4 \pi^2 c^2} \frac{1}{h^4} \frac{1}{\beta^3} \int_0^{\infty} dx \frac{x^3}{e^x - 1}$$

$\beta h \omega = x$   
 $d\omega = \frac{dx}{\beta h}$



So, we will write down  $dE$  instead of writing a delta, we will write down  $dE$ . Hence the total energy throughout the whole spectrum is going to be  $E$ , which is going to be  $\frac{h}{4 \pi^2 c^2} \int_0^{\infty} d\omega \frac{\omega^3}{e^{\beta h \omega} - 1}$ .

If you substitute  $\beta h \omega$  is equal to  $y$  or let us say  $x$ , this would be a preferred choice then you see,  $d\omega$  is going to be  $\frac{dx}{\beta h}$ . So, that you have  $\frac{h}{4 \pi^2 c^2} \int_0^{\infty} dx \frac{x^3}{\beta^3 (e^x - 1)}$  which gives you  $\frac{h}{4 \pi^2 c^2} \frac{1}{h^4} \frac{1}{\beta^3} \int_0^{\infty} dx \frac{x^3}{e^x - 1}$ . This is once again familiar to us.



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$$= \frac{k}{4\pi^2 c^2} \frac{1}{h^4} \frac{1}{\beta^3} \int_0^\infty dx \frac{x^3}{e^x - 1}$$

Stefan Boltzmann Law:

$$E = \frac{(k_B T)^4}{4\pi^2 c^2 h^3} \Gamma(4) \zeta(4) = \sigma T^4$$

$$\sigma = \frac{k_B^4}{4\pi^2 c^2 h^3} \Gamma(4) \zeta(4) = \frac{\pi^4}{60} \frac{k_B^4}{h^3 c^2} = \frac{\pi^2}{60} \frac{k_B^4}{h^3 c^2}$$

Stefan's Constant

$$dE = \frac{k}{4\pi^2 c^2} \frac{\omega^3}{e^{\beta\hbar\omega} - 1} d\omega$$



This is gamma 4 zeta 4. So, that you have this expression, the total energy, which is radiating out of this hole is going to be 1 over 4 pi square C square h bar cube k B T raised to the power 4, gamma 4, zeta 4.

And this is sigma T to the power 4, where sigma is going to be k B raised to the power 4 divided by 4 pi square, C square, h bar cube times gamma 4 zeta 4. which is going to be pi 4 over 60 and then you have k B raised to the power 4 pi square, h bar C whole cube so, that you have pi square k B raised to the power 4 divided by 60 h bar C whole no, sorry, this is wrong here. So, this is going to be h bar cube C square, h bar cube C square.

This law that you see over here is called the Stefan Boltzmann law. And this constant is called Stefan's constant. One can plug in the values of k B h bar and C to determine its value. As a concluding remark, we come back to this expression, we come back to this expression

that we have over here. And we look at; so, we have  $dE$ , which is going to be  $h$  bar over  $4\pi$  square,  $C$  square and then I have  $\omega$  cube  $e$  to the power  $\beta h$  bar  $\omega$  minus 1  $d\omega$ .

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Stefan's Constant

$$dE = \frac{h}{4\pi^2 c^2} \frac{\omega^3}{e^{\beta h \omega} - 1} d\omega \quad \beta h \omega \gg 1$$

$$dE = \frac{h}{4\pi^2 c^2} \omega^3 e^{-\beta h \omega} d\omega$$
  

$$dE = \frac{h}{4\pi^2 c^2} \frac{\omega^3}{\beta h \omega} d\omega \quad \beta h \omega \ll 1$$

$$dE = \frac{k_B T \omega^2}{4\pi^2 c^2} d\omega \quad \text{Rayleigh Jean Lo.}$$

The slide also features the NPTEL logo and a small video inset of a man in a blue shirt.

Now, if  $\beta h$  bar  $\omega$  is much much larger than 1, we have this as  $dE$ , which is going to be  $h$  bar over  $4\pi$  square  $C$  square  $\omega$  cube  $e$  to the power minus  $\beta h$  bar  $\omega$   $d\omega$ . In the opposite limit, when  $\beta h$  bar  $\omega$  is much much less than 1, I can expand the exponential over here and that gives me  $dE$  is going to be  $h$  bar over  $4\pi$  square,  $C$  square  $\omega$  cube divided by  $\beta h$  bar  $\omega$ .

To the leading order; the  $h$  bar  $h$  bar cancels out and of course, you have a  $d\omega$  over here, which is going to be  $\omega$  square over  $4\pi$  square  $C$  square times  $k_B T d\omega$   $dE$ . And this is the expression that you get, when  $\beta h$  bar  $\omega$  is going to be much much less than

1. This is what is called a, what is known as the Rayleigh Jeans law and this is called Wien's law in radiation. So, with this we conclude this section.