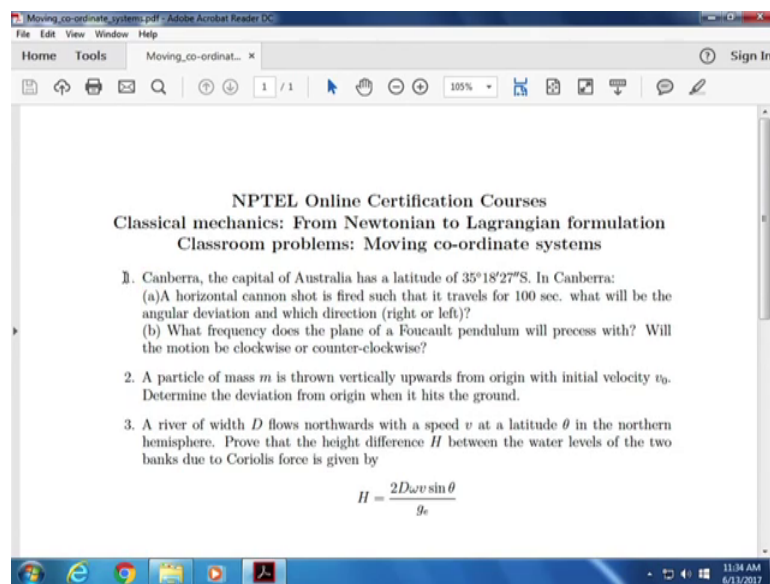


Classical Mechanics: From Newtonian to Lagrangian Formulation
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Lecture – 26
Moving Co-ordinate Systems – 4

We start this class by looking into some problems with which has vertical sorry horizontal motion involved in it.

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So, the first problem is number 1 Canberra the capital of Australia has a latitude of 35 degrees 17 minutes and 27 seconds. In Canberra, a horizontal cannon shot is fired such that it travels for 100 seconds, what will be the angular deviation and in which direction right or left.

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$$1) \quad 35^{\circ} 18' 27'' \equiv 35 + \frac{18}{60} + \frac{27}{3600} = 35.3075^{\circ} \text{ S}$$

$$= \cancel{35} + \frac{18}{60} + \frac{27}{3600}$$

$$a) \quad \alpha = \omega \sin \theta t = (7.3 \times 10^{-5} \times 0.578 \times 100) \text{ m}$$

$$= 0.42 \text{ cm}$$

$$\alpha = 0.42 \text{ rad}$$

$$b) \quad \omega_2 = \omega \sin \theta$$

$$= 7.3 \times 10^{-5} \times 0.578$$

$$T = \frac{2\pi}{\omega \sin \theta} = \frac{1 \text{ day}}{\sin \theta} = \frac{1}{0.578} \approx 1.7 \text{ day}$$

So, for this particular problem, what is given is the latitude which is 35 degrees 18 minutes and 27 seconds.

First of all what we need to do is we need to bring it to a decimal format which will be 35 plus 18 by 60 plus 27 by 3600; now if you perform this it will be something like 35 point. So 18 by 60 so, it will be I will use my; I do not have a calculator, but I have a mobile which I can use as calculator. So, if I do that it will be 35 point sorry 35.3 for this particular term and 27 divided by 3600 plus 3 yeah. So, finally, it will be 35.3075 right and it is given as south that is very important.

Now, the angular deviation alpha is given as, omega sin theta t omega is the angular velocity of earth's rotation, sin theta is the theta is a latitude which is given as 35.3075. So, all we need to do and t is given as 100 seconds. So, we have to put these values in. So, it will be seven points 3 into 10 to the power minus 5 into sin of this angle. So, once again I use my calculator here. So, it will be 0.578 into t is given as 100 seconds. So, the deviations and if we calculate this we will get this multiplied by 7.3 into 100 equal to divided by. So, it will be yeah 0.42 centimeters approximately right.

So, this is the angular or sorry 0.42, it is a dimensionless number. So, 0.42 will be the angular deviation; alpha equal to 0.42 radians this is the answer right. So, this is question number 1 a, question number 1 b is for the same place; that means, at Canberra which is given by 35.3075 then 3075 degrees south latitude, what will be the a frequency of a

plane of Foucault pendulum. So, let us come to the question here what frequency does the plane of a Foucault pendulum will process with will the motion be clockwise or counter clockwise.

So, there is a second part of part I mean second part of this. So, we have to tell with whether the deviation will be towards the right or left, now it happened to be at the southern hemisphere this s means southern hemisphere and in southern hemisphere the deviation has to be left to the original direction. So, left is the right answer right. Now let us move to the second part what is the plane, what is the frequency of the plane or does the plane of a Foucault pendulum will process with.

So, we know that the frequency is given by $\omega = \omega_z \sin \theta$, which is given by $\omega \sin \theta$ once again right. So, again we are not looking for the time period we are just looking for the frequency. So, we have to multiply 7.3×10^{-5} , which is ω that has to be multiplied with 0.578. So, this will give us some number and now if I want to calculate the time period T which will be given by $2\pi / \omega \sin \theta$. Now recall that $2\pi / \omega$ is 24 right. So, it will be $2\pi / \omega$ is 1 day essentially. So, it will be 1 day divided by $\sin \theta$, which will be 1 divided by 0.578 which will be once again I use my calculator ok.

So, 1 divided by 0.578 will be 1.78 day approximately. So, it will take almost yeah more than 1 and half days for the plane of Foucault pendulum to make a complete rotation. So, if we start here let say to begin with then the plane, again the plane will rotate whether it will rotate clockwise or anticlockwise. Now please recall in southern hemisphere for a person standing on southern hemisphere the earth seems to rotate clockwise right. So, the plane of the pendulum will rotate anticlockwise right. So, it will make a complete rotation in 1.7 days right.

Now,. So, with this we move to our next topic, which is motion under I mean a deviation due to Coriolis force under vertical motion.

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The image shows handwritten notes on a blue background. On the left, a 3D coordinate system is drawn with axes labeled \hat{e}_x (East), \hat{e}_y (North), and \hat{e}_z (Up). A vertical dashed line of height h is shown along the \hat{e}_z axis. The velocity vector is given as $\vec{v} = -v_z \hat{e}_z = -\dot{z} \hat{e}_z$. Below this, the equations of motion are derived: $\ddot{x} = 2\omega \cos\theta \dot{z}$, $\ddot{y} = 0$, and $\ddot{z} = -g \Rightarrow \dot{z} = -gt$. A boxed equation shows $z = h - \frac{1}{2}gt^2$, and the time to fall is given as $t_m = \sqrt{\frac{2h}{g}}$. On the right, the Coriolis force is calculated: $F_c = 2m \vec{v} \times \vec{\omega}$. The angular velocity vector $\vec{\omega}$ is expressed in terms of unit vectors: $\vec{\omega} = \omega \cos\theta \hat{e}_z + \omega \sin\theta \hat{e}_y$. The cross product is shown as a determinant:
$$F_c = 2m \begin{vmatrix} \hat{e}_x & \hat{e}_y & \hat{e}_z \\ 0 & 0 & -\dot{z} \\ 0 & \omega \cos\theta & \omega \sin\theta \end{vmatrix}$$
 This is then simplified to:
$$F_c = 2m \left[\dot{z} \omega \cos\theta \hat{e}_x + \hat{e}_y [0] + \hat{e}_z [0] \right] = +2m \dot{z} \omega \cos\theta \hat{e}_x$$

Let us consider this situation here we have \hat{e}_x \hat{e}_y \hat{e}_z , now if you recall \hat{e}_x is east, \hat{e}_y is north and this is up right. Now if this is the situation then if a particle is falling from rest at some height h from the horizon, then we can try modeling this by writing the velocity \vec{v} vector is given by minus v_z it will be \hat{e}_z , cap actually or we can simply write minus \dot{z} dot \hat{e}_z cap this we can do right.

Now, if we calculate the Coriolis force or yeah; Coriolis force on this particle it will be twice $m \vec{v}$ cross $\vec{\omega}$ right and once again if we break it into components it will be \hat{e}_x \hat{e}_y \hat{e}_z . Now for velocity the only component is \dot{z} dot right and for $\vec{\omega}$ the component is 0 , $\omega \cos\theta$ $\omega \sin\theta$ right. So, if we continue we will get twice $m \hat{e}_x$ caps, it will be minus \dot{z} dot $\omega \cos\theta$ right plus \hat{e}_y cap which will be given by; So, 2 terms here yeah. So, it will be given by 0 because both will cancel this column is 0 and \hat{e}_z cap which is given by once again this and this is equal to 0 . So, it will be once again equal to 0 right.

So, the only existing term will be twice $m \dot{z}$ dot $\omega \cos\theta$, there is a minus \hat{e}_x cap right. So, we have already taken a minus here oh sorry. So, the there will be a minus here, but this will be a minus here. So, this will be a plus. So, this will be a plus right. So, I can write down the equations of motion in all the direction. So, \ddot{x} double dot will be $2 \omega \cos\theta \dot{z}$ dot \ddot{y} double dot is equal to 0 and \ddot{z} double dot is equal to minus g . So, these are the equations of motion which we need to solve.

Now, for the last equation if we integrate it once we get \dot{z} is equal to minus g t right and integrating once more we get z is equal to h minus half g t square. So, now we have included the initial condition that is started from a height h . So, if you put t equal to 0 you get equal you get h z equal to h right. So, if we include that here; also one thing we can do is we can find out the total time of flight, that the particle takes from coming falling from this point to that point and that is given by t_m which is simply $2h$ by g root right. So, we will just keep this for future.

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The image shows a handwritten derivation on a blue background. On the left, a 3D coordinate system is shown with axes \hat{e}_x , \hat{e}_y , and \hat{e}_z . A particle is shown falling from a height h along the \hat{e}_z axis. The velocity vector is $\vec{v} = -\dot{z}\hat{e}_z$. The Coriolis force is calculated as $\vec{F}_c = 2m\vec{v} \times \vec{\omega}$, where $\vec{\omega} = \omega \cos\theta \hat{e}_x + \omega \sin\theta \hat{e}_z$. The resulting force is $\vec{F}_c = 2m\dot{z}(\omega \cos\theta \hat{e}_y)$. The equations of motion are $\ddot{x} = 2\omega \cos\theta \dot{z}$, $\ddot{y} = 0$, and $\ddot{z} = -g$. The time of flight is $t_m = \sqrt{2h/g}$. The final displacement in the x direction is $x_m = \frac{1}{3}\omega \cos\theta \left(\frac{2h}{g}\right)^{3/2}$, labeled as 'Dwe East'.

So, we have a particle which is moving which is falling from an height h , particle of mass m which is not very important, we will see once again we level this as e_x e_y and e_z . So, we want to see how much this particle will deviate from. So, ideally in an ideal case if it starts exactly on z axis then it should fall at the origin. So, you want to see which direction and by what amount this particle will deviate due to the Coriolis force.

Now, as the particle is falling straight, we can assume the velocity v we can write it as minus z dot e_z cap. So, this is the form of the velocity. So, the Coriolis force F_c will be $2m$ v cross ω , which is given by e_x cap e_y cap e_z cap 0 0 minus z cap. So, we have included this minus here and we have 0 ω \cos θ and ω \sin θ right. So, if we do that we see finally, only the e_x component will survive because we have a 0 column here. So, if we calculate e_x it will be this 0 minus of minus so, there will be a plus ω \cos θ z dot right.

So, if I write the force equations along 3 different axis, we get $x \ddot{} = 2\omega \cos \theta$ or m will cancel out because we are just writing the acceleration $2\omega \cos \theta$, $z \ddot{} = -g$ and $y \ddot{} = 0$. Now for the last relation if we integrate it we will get $\dot{z} = -gt$ and $z = h - \frac{1}{2}gt^2$ and from here we can calculate the maxi, the time of flight which is the time which it will take to reach the ground, if we put $z = 0$ we get t_{\max} which is $2h/g$ right.

Now, what we can do is we can substitute this with gt and please remember we cannot write minus gt here because we have already included the minus sign here. So, what we need to do is we need to write $x \ddot{} = 2\omega \cos \theta \cdot gt$, now integrating it once will give you $\omega \cos \theta \cdot g t^2$ this 2 will cancel plus some constant c_1 . Now please recall that at $t = 0$ $x = 0$ and $\dot{x} = 0$. So, $x = 0$ means the initially it starts along above z axis only above the origin on z axis only. So, there is no x or y deviation, the second assumption essentially means that although ok.

So, now think of this situation we have already calculated the Coriolis force and we have seen that Coriolis force is acting along the east direction. Please remember that north sorry e_x is sorry e_x is east e_y is north and e_z is up right. So, the Coriolis force is acting along east direction. So, there will be a deviation along east, now this deviation along our force will cost a deviation along east; that means, there will be a velocity component along east as well. Now this velocity component along east will insert another force on in the y direction. So, there will be all the secondary effects. But because this force itself is too small it is proportional to $\omega \cos \theta$. So, we are assuming that the secondary effects are vanishing non vanishing effect. So, we are setting $\dot{x} = 0$, $\dot{y} = 0$ at any t . So, this is a very important assumption we are making here and this is a valid assumption that means, we are neglecting the second order effect. So, this is a valid assumption given that the second order effects are actually very small. So, we can use that $\dot{x} = 0$ which immediately sets $c_1 = 0$.

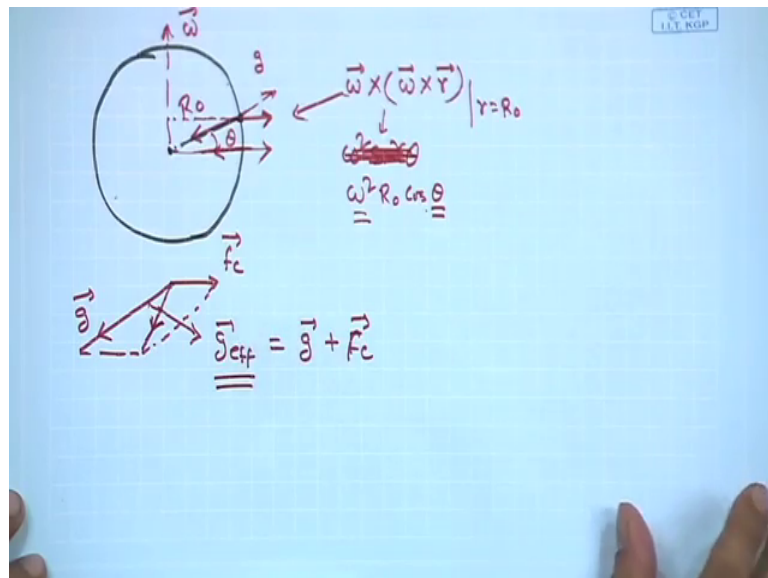
Similarly, we can integrate it once more to get $\frac{1}{6}\omega \cos \theta \cdot g t^3$ plus some c_2 which we can also set equal to 0. So, the maximum deviation is when we put $t = t_{\max}$ and we immediately see that this is nothing but $\omega \cos \theta$. Now there will be a $2h/g$ cubed by g cubed this will cancel out; oh sorry there is a root here sorry. So,

if we include this and do the calculation, the final expression will be $8h$ cubed by g yeah. So, there will be a g outside and there will be a g to the power 3 by 2 inside minus 3 by 2 so it will give you h to the power half right.

So, this is the maximum deviation and the maximum deviation is due east. So, this is the maximum deviation and this is due east right, because x axis if you recall is the eastern direction. So, this is the important result because we start from the top and we see that in the deviation is given due east and also please note that the expression, what is the final expression we got is $\omega \cos \theta$, I mean it is proportional to $\omega \cos \theta$. So, does not matter if we stay in the northern hemisphere or southern hemisphere it will give you a deviation which is due east right.

So, with this we also briefly introduced the effect of centrifugal acceleration that is the last topic of this chapter.

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So, this is our earth any point on earth which is at the radius of earth is R_0 , at any point the centrifugal acceleration which is given by $\omega \times \omega \times r$ at any point r equal to R_0 , I can you can carefully consider the components of ω and r is a radial vectors. So, the direction of r at this particular point is this. So, if you do this careful a calculation I mean if carefully check the sign of this vector, then it can be shown that this vector is just at any point it is along this particular direction.

So, basically if I join that point with the axis of rotation this will be the extended direction of this particular vector. Now what is the effect? Recall, that the force due to gravity which is which will cause g the acceleration due to gravity is along this line. So, basically g vector acts along this line, now this force is there all that is also force which is omnipresent and the magnitude of this term will be $\omega^2 r \sin^2 \theta$ being the latitude. So, I am leaving it up to you to prove this that $\omega \times \omega \times r$ will have a magnitude of this and direction of this ok.

So, this I can so, sorry it will be $\omega^2 r \cos \theta$ I am sorry. So, it will be $\omega^2 R \cos \theta$ not that right. So, if this is the case then this is a force which is present at any point of course, the magnitude will change depending on the value of θ right. So, if it is $\cos \theta$ dependence then it will be 0 at the poles and maximum at the equator right. Now at equator does not have much of a problem because it for this one acts along this line this one the g acts along that line. So, they kind of cancel each other.

So, but at any point they are not in the equal direction, I mean equal or opposite sorry they are not in the opposite direction. So, the effective g at any point if this is my direction of g vector, this is my direction of this particular centrifugal acceleration which we will call f_c . So, my effective g is a vector sum of this 2 right. So, this is my g effective which is a vector sum of g plus f_c right. So, this is not. So, the g effective we as we see is not along the line joining that particular point to the center of the earth, but which is a line which is slightly deviated from it; of course, this deviation will be small because it has a ω^2 term the this f_c has ω^2 term, but it is not 0 there is a considerable amount of deviation and this particular line is called the plumb line ok.

So, anyway this is what I wanted to express briefly. So, all I can tell you that, you can calculate in principle the change in magnitude of g at any point on earth, using this relation right. Now with this we move to the second problem of this problem, set the particle of mass m is thrown vertically upwards from origin with initial velocity v_0 determine the deviation from origin when it hits the ground, now it is a slight deviation of the problem which we have already performed instead of starting from a height h it is starting from the origin let we say.

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$\hat{e}_z (U)$
 $\hat{e}_y (N)$
 $\hat{e}_x (E)$

$v = v_0 - gt$
 $\ddot{z} = -g \Rightarrow \dot{z} = v_0 - gt$
 $\ddot{x} = -2\omega \cos\theta \dot{z}$
 $\ddot{x} = -2\omega \cos\theta (v_0 - gt)$
 $x = -2\omega \cos\theta \left(\frac{v_0 t^2}{2} - \frac{1}{6} g t^3 \right)$
 $x_m = -\frac{4}{3} \omega \cos\theta \frac{v_0^3}{g^2}$
 Due west

$z = v_0 t - \frac{1}{2} g t^2$
 $t_m = \frac{2v_0}{g}$

So, we have \hat{e}_x , \hat{e}_y , and \hat{e}_z cap. So, the particle of interest starts from this in a with a initial velocity v_0 in the vertical direction and then reaches a maximum height and then it comes back here right. So, once again we have to write the equation of motion, here please understand that v is given by $v_0 - gt$ right. So, I will not write a vector equation that will be bit confusing, the magnitude of v is at any point is given by $v_0 - gt$ right. So, as the velocity is in the upward direction what we can do is, we can once again calculate the coriolis forces and it can be shown I am leaving it to you for to do the exact calculation, this will be $2\omega \cos\theta \dot{z}$ right \ddot{x} as usual is equal to 0 and \ddot{z} is equal to minus g .

Now, we see that because the motion is in the upward direction there is a negative sign. So, the force is not towards east, but towards west right. This is up now if we again substitute \dot{z} which will be given by $v_0 - gt$ at any point keep taking this initial consideration into account. So, we get $\ddot{x} = 2\omega \cos\theta (v_0 - gt)$ right. So, after once again we have to integrate it twice and the final value will be I am just writing the final expression will be minus $2\omega \cos\theta v_0 t^2$ by $2 - \frac{1}{6} g t^3$, right. So, this will be the final expression.

Now, z if I integrate we will get $v_0 t - \frac{1}{2} g t^2$ right, now from here what we can get is once again a total time of flight by setting z equal to 0 it is a quadratic. So, it will have 2 roots 1 for t equal to 0 will have z equal to 0 and once for $2v_0/g$ just directly from this relation z equal to 0 and you will get $t_m = 2v_0/g$. So, when we put this into this relation to get a total maximum deviation, what we get is

minus 4 by 3rd omega cos theta v 0 cubed by g square right, so this is the final answer and because of the minus sign this deviation is due west.

Now this is a very interesting observation because if you recall in the previous problem what we did, the deviation was due east and now the deviation is due west. Now in the previous problem we started from this point and the particle came hit the ground not here, but somewhere along x axis right, but in this case when we start from here the particle comes back and hits the ground somewhere in the y, somewhere along minus x axis. Why it is so I am just leaving on to you.

Now, before we finish what we can do is we can quickly look up into this problem once more, look at this particular expression for xm which was the case when the particle was dropped from a certain height. So, by the way if you find out the maximum height h and reduce this expression in terms of that, it will be exactly you will find that it will be exactly 4 times this expression there with the minus sign here.

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Handwritten notes on a whiteboard showing the calculation of horizontal deviation x_m for two different launch angles. The first case is for $\theta = 45^\circ$, resulting in $x_m \approx 1.54 \text{ m (E)}$. The second case is for $\theta = 90^\circ$, resulting in $x_m \approx 8.8 \text{ m (W)}$.

$$x_m = \frac{1}{3} \omega \cos \theta \left(\frac{8h^3}{g} \right)^{1/2}$$

$h = 100 \text{ m}, \theta = 45^\circ$ | $\theta = 90^\circ$
 $x_m \approx 1.54 \text{ m (E)}$ | $x_m \approx 2.2 \text{ m (E)}$

$$x_m = -\frac{4}{3} \omega \cos \theta \left(\frac{8h^3}{g} \right)^{1/2}$$

$x_m \approx 16 \text{ m (W)}$ | $x_m \approx 8.8 \text{ m (W)}$

So, let us examine this expression then we will get a field for this expression also, very quickly x_m is equal to one third omega cos theta 8 h cubed by g to the power half.

So, if we put say x is equal to 100 meters sorry h is equal to 100 meters, what are the values of this deviation we get. So, for theta equal to 45 degrees, does not matter if it is north or south because it is a cos theta dependence, we will get x_m is approximately 1.54

centimeters and for the same height θ equal to 90 degree, means we will get sorry θ equal to 0 degree, where we will get maximum deviation we will get x is equal to x_m is equal to 2.2 centimeters.

So, you see that the deviations are small when we are compared to a height of 100 meters the deviation is small, but now consider the case when the object will be deviated from a height of few kilometers, then these numbers are not small so these are significant numbers. Now as I said when we are throwing the object this particular case, when we are throwing the object from this one from this point and it comes back this it is not only due west, but it will be exactly the same expression multiplied by 4. So, in this particular case x_m will be minus 4 by 3rd ω cos θ 8 h cubed by g to the power half.

We have to substitute, we have to find and we have to replace v_0 by h . So, in that case if we use the same number x_m will be so this will be towards east. So, it will be almost 6 centimeters towards west and in this case it will be x_m approximately 8.8, approximately 9 centimeter towards west. So, you see these numbers are not very insignificant anymore. So, if we can throw a stone at up to a height of 100 meters, when it comes back it deviates by 6 centimeters at 45 degree latitude and that 0 degree latitude the deviation is almost 9 centimeter. So, these numbers are not very insignificant with this we end this class. So, we will see you in the next class.

Thank you.