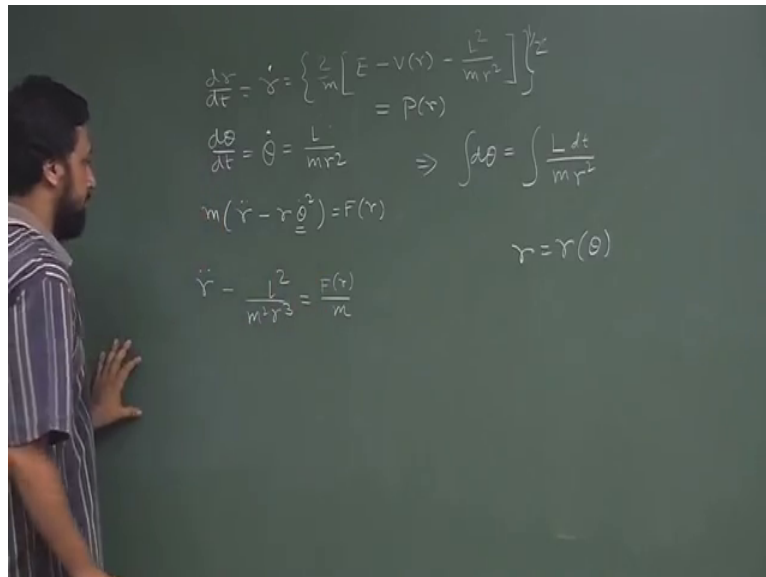


Classical Mechanics: From Newtonian to Lagrangian Formulation
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Lecture – 10
Central forces – 3

So, in the last class, we have seen that for a motion under central orbit. We can write \dot{r} equal to $\sqrt{2/m[E - V(r) - L^2/2mr^2]}$, which might not be seeing easily, but anyway you know this expression. And $\dot{\theta}$ equal to L/mr^2 .

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Now this one, we can write this as total derivative of r and t . And from here we can just shift this in order to write \dot{r} as $\sqrt{2/m[E - V(r) - L^2/2mr^2]}$. So, instead of writing this full expression we can just replace it, we can just call it some P of r , P being some function now in this particular function and because it has r here and here. So, we can just write this as $dr/P(r)$ this integration and dt integration.

Now, performing this integration, will give me r as a function of time. Right now this integration on the sorry this equation for the we for what we have for $\dot{\theta}$ is $\dot{\theta} = L/mr^2$. So, this can be rearranged to get integration $d\theta = L dt/mr^2$.

So, once we solved, the once we could integrate this part, we can take r as a function of time. And then we can substitute it back in this equation and then we can get θ as a function of time. So, essentially we get r as a function of time, θ as a function of time. And in principle once we can perform these 2 integrals, the problem is totally solved. So, we know that we know r and θ of the entire trajectory as a function of time which essentially it is equivalent of saying that we know the entire trajectory as a function of time. That is the total solution of the problem we are looking for.

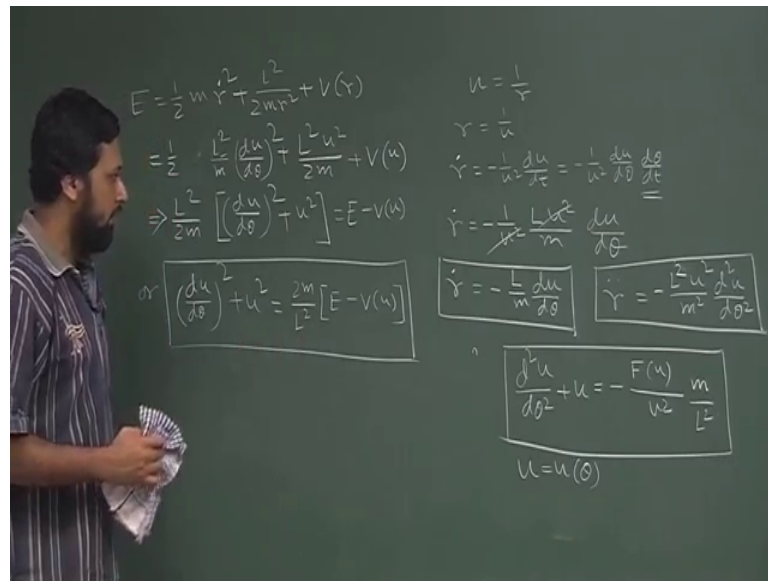
Now, there is a technical issue. Solving this equation is hard. Because rather integrating this expression in this particular form is hard. Because \Pr first of all it is not a very simple looking expression. Secondly, this v_r here is there is a a to the power half here. I think now you can see it yes right it is fine. So, V of r is not a known function. So, we do not do not know it might have singular it is a singularity at some point. It could diverge at very low or very high values of r .

So, we do not want to get the we do not want to perform this integral in a sense that it might not be a very easy task to perform especially by hand. So, in order to gain insight, we can also try to get solution of the form r does a function of θ . This will be a time independent form of the orbit. So, this will also give you the orbit, but we does not include the time explicitly. So; that means, it will give you the orbit not as a function of time.

But it will give you the orbit as a functional form of r as a function of θ . Now in order to do that order to get this solution, we need to get an equation, differential equation which includes only r and θ and there is no t dependence. How we can do that? We can do that by start writing recalling that $m r \ddot{\theta} - r \dot{\theta}^2$ is equal to F_r . So, this is the standard equation of portion of central force under the central force F_r . Now if we replace $\dot{\theta}$ by this expression, this reduces to $r \ddot{\theta} - r L^2$ square by $m^2 r$ to the power 4 is equal to F_r by m or just simplify a bit get rid of this one. So, it will be simply r^3 here.

Now, what do we do? We introduce a new set of parameter which we call u as $1/r$.

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So, it is just the inverse of the radial part. So, if this is the case. Then r is also one over u . \dot{r} is minus 1 by u square du/dt , we can do a mathematical manipulation on this. And we can write minus 1 by u square $du/d\theta$, which is $\dot{\theta}$ again. And in this particular notation $\dot{\theta}$ which is L by $m r$ square, will be $L u$ square by m right. So, if we substitute this, \dot{r} will be minus 1 by u square. And this is your $\dot{\theta}$ $d\theta/dt$.

Which is $L u$ square by $m du/d\theta$ right. So, you see that you will cancel out leaving behind \dot{r} is equal to minus L by $m du/d\theta$. This is one relation. We can do this derivative a second time. I am leaving it to you. So, essentially you have to evaluate \ddot{r} also because this equation has \ddot{r} . So, you have to take another time derivative of this. So, just the same it will be $d^2 u/d\theta^2$ if you take the time derivative of this you can replace this, as $d^2 u/d\theta^2$ and then there is another $d\theta/dt$, which is a $\dot{\theta}$ which has to be replaced back. Once again now if you do that, your final expression I am just leaving it to you as an exercise. I think you can do that.

So, your final expression for \ddot{r} will be minus $L^2 u^2$ by $m^2 d^2 u/d\theta^2$, minus $L^2 u^2$ by $m^2 d^2 u/d\theta^2$ that is the expression for \ddot{r} . Now it is time to do the substitution here, take that substitute it here. So, it will be and also you have to substitute r for u or u for r here. So, this will be minus L

square u^2 by $m^2 d^2 u / d\theta^2$, 2 minus $L^2 u^3$ by m^2 equal to $F r$ by m .

So, immediately we see that we can get rid of m^2 or whether we can get rid of m^2 from this side. So, multiply this by m^2 , which will be just multiplying it or rather it is by m here m^2 multiply it by m^2 . So, if you take $L^2 u^2$ common, with a negative sign. And simplify this expression slightly it will be $d^2 u / d\theta^2$ plus u equal to. So, this will be $m F / L^2 u^2$ minus f . So, this will be F of u $F u$, by $F u$ by $u^2 m$ by L^2 . So, F a F / L^2 yeah right. So, it is sorry (Refer Time: 10:03) it is not look nice yeah. So, it is $d^2 u / d\theta^2$, plus u equal to minus $F u$ by $u^2 m$ by L^2 . Now in this equation, we have we can solve this equation to get u as a function of θ which is equivalent of get getting r , as a function of θ right.

So, because we already know that u and r they are just related by $u = 1/r$, this is the equation we will be using mostly for our description of centre I mean getting in description of central orbit or sometime we will be using it to find out law of force, sometime we will be using it to get the differential equation, when we know the law of force we will take up this examples very soon and we will see. Now before going there we can also get in get an alternative equation which is not a homogeneous equation, but that is a first order equation. If you recall the energy expression was E equal to kinetic energy plus potential energy. So, if you do that you will simply get $\frac{1}{2} m \dot{r}^2$ plus $L^2 / 2 m r^2$ plus $V(r)$ right. So, if we now substitute \dot{r} with this expression and r with u then energy expression becomes $\frac{1}{2} m L^2 d u / d\theta^2$ plus $L^2 u^2$ by twice m plus b of u right.

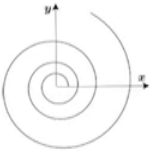
Now, rearrangement of this gives you. So, you see one m cancels immediately. And we have we can take $L^2 / 2 m$ common right. If we take $L^2 / 2 m$ common, we have $du / d\theta^2$ whole square. So, I am just writing in your step here u^2 is equal to u minus $V(u)$ or sorry not this, $du / d\theta^2$ whole square plus u^2 equal to $2 m$ by $L^2 u$ minus. So, this is another equation which is not exactly. So, this is the second order homogeneous differential equation. This is the first order in homogeneous differential equation because we have a du / dt whole square.

But this is also a this is also usable at some point it is also useful at for certain ex certain cases we will see that. So, now before moving into any further moving any further in the by using this equation and see what happens for different cases, let us take some examples. So, our have we have some classroom problems like we have in our every class.

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Classical mechanics: From Newtonian to Lagrangian formulation
Classroom problems: Central Forces-part 1

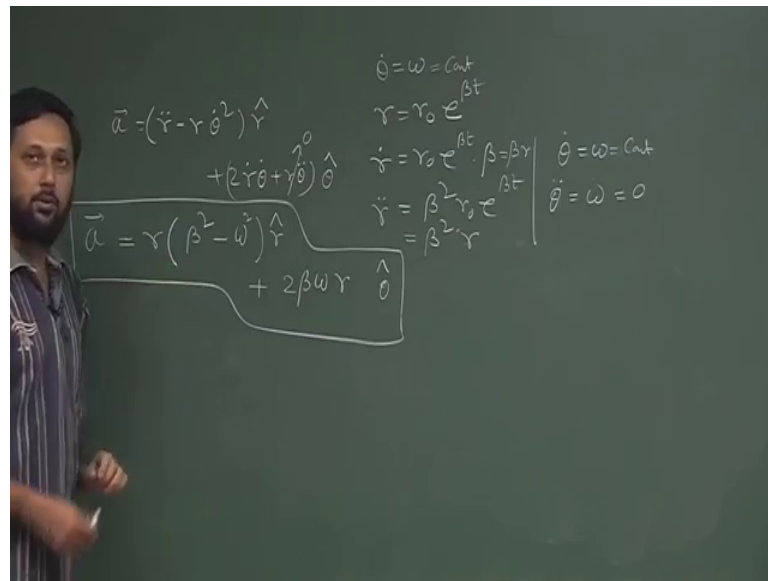
1. A particle moves along a spiral described by the equation $\dot{\theta} = \omega = \text{constant}$, $r = r_0 e^{\beta t}$. Find an expression for acceleration \vec{a} .



2. Prove that the parametric equation $r = -at$, $\theta = -b/t$ where a, b constants, represents a central orbit. Find the force law.
3. For the particle's trajectory given in problem 1, set $\beta = \omega$ and find the law of force $F(r)$.
4. A particle of mass m moves under the influence of a central force given by $F(r) = \frac{2L^2}{m^2 r^3}$. Set up the differential equation of the orbit.

So, the first problem is a particle moves along a spiral described by the equation $\dot{\theta} = \omega = \text{constant}$, $r = r_0 e^{\beta t}$. We have to find an explicit expression for the acceleration \vec{a} . So, what is given it is given that $\dot{\theta}$ is equal to ω which is equal to a constant $r = r_0 e^{\beta t}$ right?

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So, if this is r_0 r dot will be $r_0 E$ to the power βt times β , double dot will be $\beta^2 r_0 E$ to the power βt right. Similarly, θ dot is equal to ω equal to C_1 constant θ double dot is equal to $\dot{\omega}$ equals to 0. So, the expression for a in order to write an expression for a we have to keep in mind that it has both radial and transverse component.

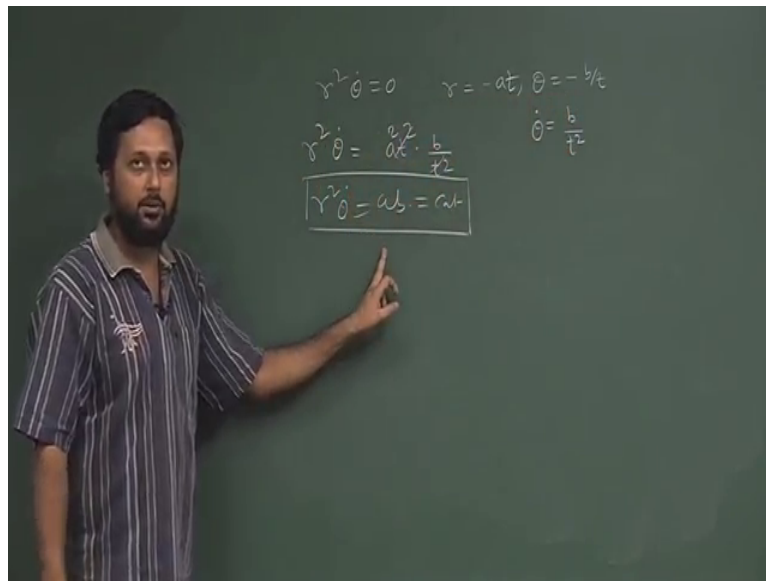
So, we have a is equal to r double dot minus r θ dot square r cap, plus $2 r$ dot θ dot r θ double dot θ cap right now a double dot right. So, if we perform this first I mean of if, we if we expand this first part and put values from sorry no not expand. If we just put values here. So, we get. So, this is essentially $\beta^2 r$. So, we get r out have $\beta^2 r$ minus $\omega^2 r$ plus. Now θ double dot is equal to 0. So, this term does not contribute to our final expression r dot is $r_0 E$ to the power βt times β , which is essentially βr . So, we have $2\beta r$ or I will just write it down for clarity plus $2\beta r$ into $\omega \theta$ dot right.

So, all final we will just rearrange the second term bring the constants $\omega r \theta$ dot. So, we have this as our answer. So, we have r times β^2 minus $\omega^2 r$ equal r cap plus $2\beta \omega r \theta$ cap. So, this is the final expression for the acceleration a in this particular case good. Now let us move to the second problem which says find the parametric equation r equal to $\frac{a}{b^2} (1 - \cos \theta)$ where a and b are constant. We prove that the parametric equation represents the central orbit and

also we have to find the force law right. Now in order to find out the weather if it is a central orbit for doing it do let us come to the board again.

So, we have to first prove. Now what is the central orbit central orbit means $r^2 \dot{\theta}$ will be equal to 0. So, if we can prove that for this particular set of parametric equation which is r equal to minus a theta equal to minus b by t , we can prove that $r^2 \dot{\theta}$ is equal to 0.

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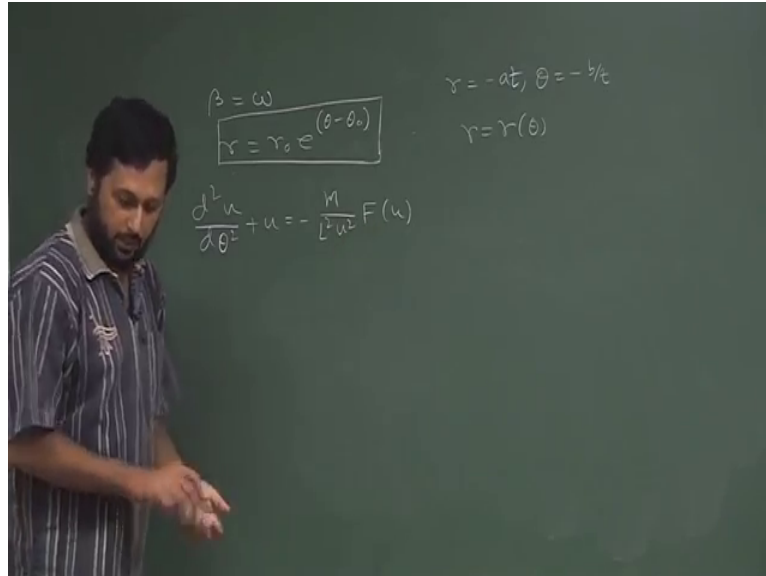


Then that essentially proves that this particular set of parametric equation represents a central orbit. Now theta is equal to minus b by t. So, theta dot is equal to b by t square t being t be a time. So, in this case we can think of t as the time right. So, this is the parameter time is the parameter, in this parametric equation. So, essentially the equation of the orbit is given in terms of a the terms of the parametric equations all we need to do is we need to eliminate. So, in order to we will come to that later.

So, we first estimate theta dot. So, $r^2 \dot{\theta}$ will be nothing, but minus a t s sorry minus will not be there, a square t square into b by t square which will be ab. So, $r^2 \dot{\theta}$ is equal to ab equal to a constant. So, that is our equation. So, we essentially proved that this orbit represented by this particular set of parametric equation is a central orbit. Now the second part is we have to find out right. Now we have to find out the equation of the orbit i t, I am living it on you because what you need to do is we need to

eliminate r . And theta I am sorry, from the expression of r and theta you need to eliminate t .

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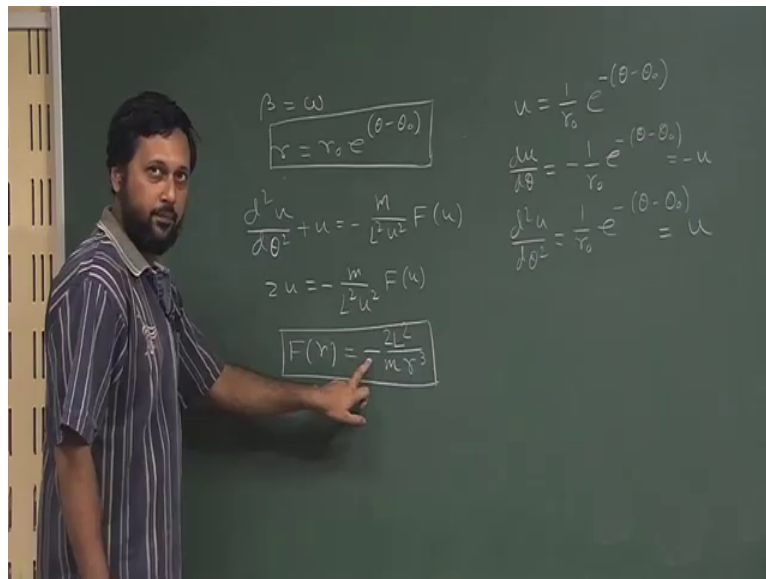


So, essentially you have to get r as a function of theta you can do that. So, I leave it to you and we moves to the x problem. So, next problem is also related to the first problem which says for the particle of that for the particle trajectory given in problem one, which is the first problem here. This spiral said v_{θ} equal to ω and find the law of force. Now this is once again this is also easy what happens is, if we set β equal to ω in this equation if you recall right r is equal to $r_0 E$ to the power βt . Now it will be $r_0 E$ to be power ωt and $\dot{\theta}$ is equal to ω equal to a constant. Now if we integrate this equation.

We get $\theta - \theta_0$ is equal to ωt . So, we substitute this for ωt . And we immediately see this equation becomes $\theta - \theta_0$. So, this is the equation of the trajectory right. So, in case when θ is equal to ωt this is the equation of the trajectory. So, we get r as a function of theta. Now what we need to do is we need to find out the law of force which is F of r . Now recall that $\frac{d^2 u}{d\theta^2} + u = -\frac{m}{L^2 \omega^2} F(u)$. So, what we need to do is we have r as a function of theta. So, we have to construct you first then we have to determine $\frac{d^2 u}{d\theta^2}$ substitute it here. And we will essentially get an expression for u F of u in terms of theta.

Now, if r is equal to this.

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So, u will be equal to $1/r_0 e^{-(\theta - \theta_0)}$. So, $du/d\theta$ is equal to $-1/r_0 e^{-(\theta - \theta_0)}$. $d^2u/d\theta^2$ is equal to $1/r_0 e^{-(\theta - \theta_0)}$, which is equal to u . So, you see that this is $du/d\theta$ oh sorry $d^2u/d\theta^2$ is essentially equal to u . And this is equal to $-u$ which is not very uncommon for exponential functions on a subsequent derivative say some time repeat themselves, sometime repeat them with a negative sign this is exactly what is happening here. So, this part is nothing, but to u and the right hand side we have $m/L^2 u^2 F(u)$ you have to bring this u^2 to this side.

And rearrangement will tell you that $F(u)$ is equal to $-L^2/m$. So, there will be a 2 here, it will be u^3 we are writing $1/r^3$. So, this is the final answer. So, we get a force law which is attractive because there is a minus sign and which is proportional to the inverse of r^3 rather $1/r^3$. So, the force which gives this type of a trajectory is an inverse cube force attractive. So, just for your curiosity the particular type of spiral which is described in this problem, when β is equal to ω this particular spiral is called a logarithmic spiral. And this is the very interesting trajectory we probably will come back to it and do some more problems.

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2. Prove that the parametric equation $r = -at, \theta = -b/t$ where a, b constants, represents a central orbit. Find the force law.

3. For the particle's trajectory given in problem 1, set $\beta = \omega$ and find the law of force $F(r)$.

4. A particle of mass m moves under the influence of a central force given by $F(r) = \frac{2L^2}{m^2 r^5}$. Set up the differential equation of the orbit.

5. A particle P has two constant velocities u and v such that u is always parallel to a fixed direction OX and v is always perpendicular to the radius vector OP.

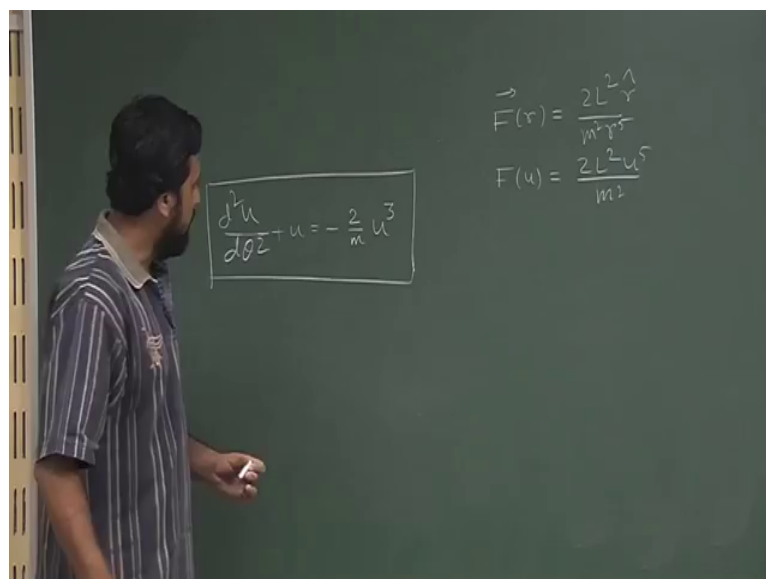
(i) Write down the velocity expressions for v , and v_θ .

(ii) Set up the equation of motion.

(iii) show that the locus of P is a conic section.

Now, let us move to the fourth problem fourth problem is a particle of mass m moves under the influence of central force given by this. So, this is a repulsive central force because there is no negative sign. And we have r to the power 5 dependence we have to set up the differential equation of this particular orbit, and the differential equation is the same $d^2 u + u = -\frac{2L^2}{m^2} u^3$.

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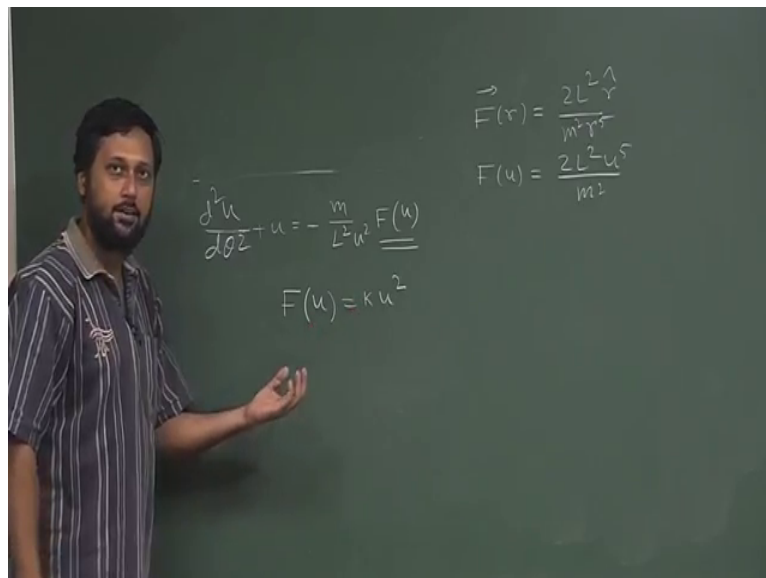


Now, what is given is the force law. So, this side is given F of r is equal to $2L^2$ square r kept by m^2 square r to the power 5 m^2 square r to the power 5 right. So, F of u .

So, we just have to replace r with $1/u$ and we immediately see that this gives you E to the power 5 square. When we replace it back here for F_u we get $2L$ to the power 5 sorry L square sorry, m square L square gets cancelled out u square is here. So, we get u cubed and also we can get rid of one m here. So, essentially the differential equation becomes $\frac{d^2u}{d\theta^2} + u = -\frac{2L^2}{m^2} F(u)$. So, this is our answer. So, all the problems what we have described here, these are very simple problem in reality sometimes we have to solve this differential equation, we will take up some more examples as we progress as we go on in our classes.

But right now, I just wanted to show you how to use this particular form of differential equation or particular form of expression. What we have got for connecting u and θ in order to first we showed how to solve this equation to get the equation of an orbit. And secondly, we sort took a force law and showed you how to set up the equation of motion. Now how to solve this equation that is a different question altogether, but of course, I took easier examples in in order to demonstrate it easily in this classroom oh we will have little more complicated problems. Next class onwards also what we are going to do is in the next class.

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We need to again recall that the original equation has a particular form, which is $m L$ square u square. Sorry I always forget this particular form and I have to look back to my notes. So, I hope you will excuse me.

So, what we will do is we will try to solve this equation for a very particular type of force which is K/r^2 . So, K/r^2 essentially means F_r is equal to K/r^2 . It could be a negative sign for attraction. There could be a positive sign for repulsion, but this force law is extremely familiar to all of us, this is the inverse square force field which consists of gravitational force which consists of electrostatic attraction or repulsion. So, this is a very common force which is which is way out which is seen universally you know in the world we live. So, we will take this up in the next class, I will see what type of particular answer do we get any very specific information about the orbit by putting this value in this particular equation.

Thank you.