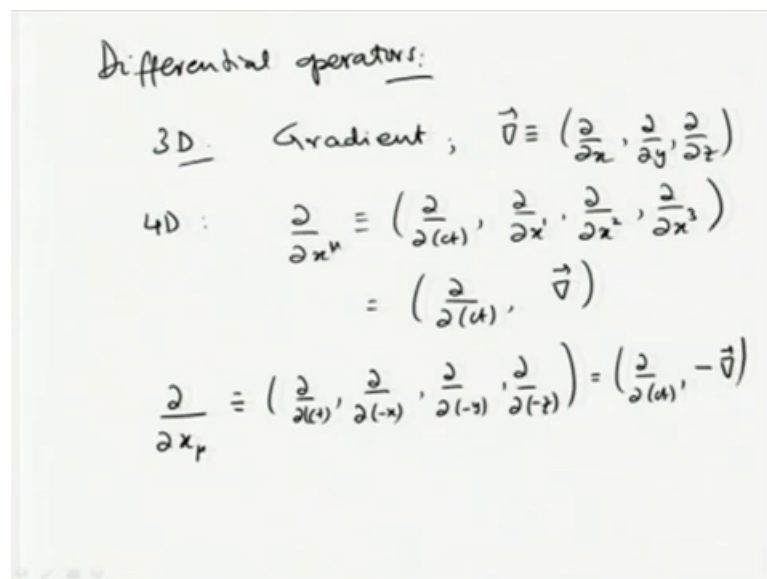


**Nuclear and Particle Physics**  
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**Module – 09**  
**Electroweak Interactions**  
**Lecture – 02**  
**Relativistic Quantum Mechanics- continued**

We will continue our discussion on the relativistic quantum mechanics. We were talking about the relativistic notations and some of the how to combine the time coordinate, time parameter and the special coordinates in a single entity called a 4 vector, and how they transform under coordinate transformations.

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Differential operators:

3D: Gradient ;  $\vec{\nabla} \equiv \left( \frac{\partial}{\partial x}, \frac{\partial}{\partial y}, \frac{\partial}{\partial z} \right)$

4D :  $\frac{\partial}{\partial x^\mu} = \left( \frac{\partial}{\partial (ct)}, \frac{\partial}{\partial x^1}, \frac{\partial}{\partial x^2}, \frac{\partial}{\partial x^3} \right)$   
 $= \left( \frac{\partial}{\partial (ct)}, \vec{\nabla} \right)$

$\frac{\partial}{\partial x_\mu} = \left( \frac{\partial}{\partial (ct)}, \frac{\partial}{\partial (-x)}, \frac{\partial}{\partial (-y)}, \frac{\partial}{\partial (-z)} \right) = \left( \frac{\partial}{\partial (ct)}, -\vec{\nabla} \right)$

Let us start with differential operators. So, let us consider differential operators in relativistic notation see in 3 dimensional relative 3 dimensional case. We have gradient defined as  $\frac{\partial}{\partial x}$  along the x axis  $\frac{\partial}{\partial y}$  along the y axis and  $\frac{\partial}{\partial z}$  along the z axis.

And the 4 dimensional generalization of this can be written as  $\frac{\partial}{\partial x^\mu}$ , which is  $\frac{\partial}{\partial ct}$  which is our 0 th component and  $\frac{\partial}{\partial x^1}$   $\frac{\partial}{\partial x^2}$   $\frac{\partial}{\partial x^3}$  which is  $\frac{\partial}{\partial x}$   $\frac{\partial}{\partial y}$   $\frac{\partial}{\partial z}$  etcetera or I

can write it in a way which will separate the time component and the spatial component returned as the gradient.

This is one way of doing it and the other is with mu as a subscript on the x which is basically a x mu covariant. So, if you take the derivative with respect to the covariant x then you will have  $\frac{\partial}{\partial x^\mu}$  because the special components of x mu covariant is minus x minus y minus z and this is essentially  $\frac{\partial}{\partial x^\mu}$  minus gradient.

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$$\begin{aligned} \frac{\partial}{\partial x^\mu} \text{ transformation} \\ x^\mu \rightarrow x'^\mu = a^\mu_\nu x^\nu \\ x_\nu \rightarrow x'_\nu = a_\nu^\rho x_\rho \end{aligned} \quad \left| \quad \begin{aligned} \frac{\partial}{\partial x^\mu} &\rightarrow \frac{\partial}{\partial (x'^\nu)} = \frac{\partial}{\partial (a^\nu_\rho x^\rho)} \\ &= \left( \frac{1}{a^\nu_\rho} \right) \cdot \frac{\partial}{\partial x^\rho} \\ &= (a^{-1})^\nu_\rho \cdot \frac{\partial}{\partial x^\rho} \end{aligned} \right.$$

$$\frac{\partial}{\partial x^\mu} \text{ is a covariant vector} \Leftarrow \underline{\underline{= a_\nu^\rho \cdot \frac{\partial}{\partial x'^\rho}}}$$

Let us look at how the  $\frac{\partial}{\partial x^\mu}$  transform and the coordinate reservation. We know  $x^\mu$  goes to  $x'^\mu$ , which is  $a^\mu_\nu x^\nu$  and  $x_\nu$  covariant goes to  $x'_\nu$ , which is  $a_\nu^\rho x_\rho$ .  $\frac{\partial}{\partial x^\mu}$  contravariant, will go to  $\frac{\partial}{\partial x'^\mu}$ , which is equal to  $\frac{\partial}{\partial x^\nu} \frac{\partial x'^\nu}{\partial x'^\mu}$ , which is equal to  $\frac{\partial}{\partial x^\nu} a^\nu_\mu$ . I can write it as  $\frac{1}{a^\mu_\nu} \frac{\partial}{\partial x^\nu}$  and this one over  $a^\mu_\nu$  is essentially  $(a^{-1})^\mu_\nu$  the component of the a inverse matrix times  $\frac{\partial}{\partial x^\nu}$ , but we saw in earlier lectures that  $a^{-1}{}^\mu_\nu$  is essentially equal to  $a_\nu^\mu \frac{\partial x^\mu}{\partial x'^\nu}$ . This is basically the way a covariant vector transform from our previous lecture. So, previous lecture says that this is the way a covariant vector transform.

So, this says that  $\frac{\partial}{\partial x^\mu}$  is a covariant vector it transform like a covariant vector. So, we could consider that as a covariant.

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Similarly

$$\frac{\partial}{\partial x_\mu} \rightarrow \frac{\partial}{\partial (a_\mu^\nu x_\nu)} = \frac{\partial}{\partial (a_\mu^\nu x_\nu)} = \left( \frac{1}{a_\mu^\nu} \right) \frac{\partial}{\partial x_\nu} = a^\mu_\nu \frac{\partial}{\partial x_\nu}$$

$\frac{\partial}{\partial x_\mu}$  : Contravariant vector

Similarly, if you consider  $\partial_\mu$  by  $\partial_\mu$  by  $\partial_\mu x^\nu$  covariant goes to  $\partial_\mu x^\nu$  prime covariant equal to  $\partial_\mu y^\nu$   $a^\mu_\nu x^\nu$  is essentially equal to  $a^\mu_\nu \partial_\mu$  by let me write it here, taking same argument as earlier this is one over  $a^\mu_\nu \partial_\mu$  by  $\partial_\mu x^\nu$ , which is essentially  $a^\mu_\nu \partial_\mu$  by  $\partial_\mu x^\nu$  now the subscript. And this is how contravariant vector transforms. So,  $\partial_\mu$  by  $\partial_\mu x^\nu$  is a contravariant vector.

So, we will actually denote these 2 in a slightly simpler notation.

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Covariant derivative

$$\partial_\mu \equiv \frac{\partial}{\partial (x^\mu)} = \left( \frac{\partial}{\partial ct}, +\vec{\nabla} \right)$$

Contravariant derivative

$$\partial^\mu \equiv \frac{\partial}{\partial (x_\mu)} = \left( \frac{\partial}{\partial ct}, -\vec{\nabla} \right)$$

Compared to coordinates,

$$x^\mu = (ct, \vec{x}) ; \quad x_\mu = (ct, -\vec{x})$$

Covariant derivative operator  $\partial_\mu$  is defined as the derivative with respect to the contravariant vector. And contravariant derivative operator is  $\partial^\mu$  all right. These are the notations that we will use when you have the subscript on  $\partial_\mu$  this is a covariant vector like the earlier notations, but then this corresponds to derivative with respect to a contravariant  $x^\mu$ .

And similarly when we say contravariant derivative it is a superscript on though the partial derivative operator, which is equal to the partial derivative with respect to the covariant  $x_\mu$ . And this is essentially equal to  $\partial_\mu$  by  $\partial_\mu c^\mu$  and minus here you know how actually a  $c^\mu$  that it is plus gradient operator  $\partial_\mu$  by  $\partial_\mu c^\mu$  minus the gradient operator.

So, here one thing to notice is that compared to the coordinate  $x^\mu$  is equal to  $ct$  and  $x^\mu$  contravariant covariant is  $ct$  minus  $x^\mu$  whereas, in the case of derivative operators the covariant vector has a plus sign in front of the gradient operator the special derivative part, and the contravariant derivative has minus sign with respect to the time component and this is exactly the opposite in the case of coordinates  $x^\mu$  and  $x_\mu$  covariant and contravariant.

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Come back to KG eqn,

$$\left( \frac{c^2 \hbar^2}{2(u)^2} - \hbar^2 c^2 D^2 + m^2 c^4 \right) \psi = 0$$

$$c^2 \hbar^2 \left( \frac{\partial^2}{\partial x^2} - \frac{\partial^2}{\partial t^2} - \frac{\partial^2}{\partial y^2} - \frac{\partial^2}{\partial z^2} \right) \psi + m^2 c^4 \psi = 0$$

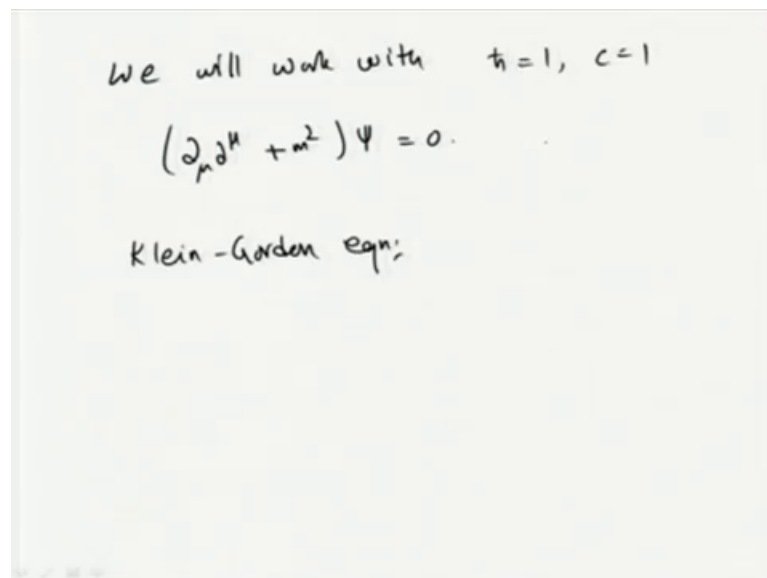
$$c^2 \hbar^2 (\partial_\mu \partial^\mu) \psi + m^2 c^4 \psi = 0$$

Now let us come back to Maxwells sorry come back to the Klein Gordon equation let me call this for short KG equation. So, we have minus  $\hbar^2$ . So, let me write it in terms of  $c^2 \hbar^2$  second derivative with respect to time or  $ct$  minus  $\hbar^2$  cross square

$c^2 \nabla^2 \psi + m^2 c^4 \psi = 0$  please refer to the previous lecture for this we have we had written down this.

So, now I can take  $c^2 \nabla^2$  as a common factor and then that will give you the square over though  $\psi$ , which is essentially let me write it as  $\nabla^2 \psi = 0$  the 0th current minus  $\nabla^2 \psi = 0$ ,  $y^2 \psi - \nabla^2 \psi = 0$ ,  $z^2 \psi$  ok,  $\psi + m^2 c^4 \psi = 0$  this can be written as  $c^2 \nabla^2 \psi + m^2 c^4 \psi = 0$  the usual dot product between 2 vectors these are all vector operators, but dot product will give you similar expression as in the case of vector,  $\psi + m^2 c^4 \psi = 0$ . Now we will work in units of  $c^2$  is equal to  $\nabla^2$  is equal to 1 ok.

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we will work with  $\hbar = 1, c = 1$

$$(\partial_\mu^2 + m^2) \psi = 0.$$

Klein-Gordon eqn;

We will work in a unit system  $\hbar$  equal to one  $c$  equal to one  $\nabla^2$  equal to 1. This is all right we can do that this is a unit system that we are considering not the SI unit Lorenz heavy said an energy it is called. And in this unit system we will have some simple this is basically to simplify our notation, because in relativistic quantum mechanics both  $\hbar$  and  $c$  will come appear in different terms. And then we if we put it equal to 1 we do not know how to write it, but then the question is whether information is lost or whether we will get confused at any point with the dimensions or with the physical quantities etcetera.

We will see that consistently at any point whenever we want to actually recover the  $\hbar$  cross and  $c$  were to put that across and  $c$  dimensional analysis will let us put these in a unique, way maybe we will take it take up some examples to illustrate this later towards the end of this lecture also, but at the moment we will take it as that we can without creating any confusion we can take  $\hbar$  cross to be equal to one and  $c$  equal to one and then work in this unit ah.

So, in that case you have  $\square \mu \square \mu + m^2$  is equal to acting on  $\psi$  is equal to 0 is our Klein Gordon equation ah.

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$$\begin{aligned} \frac{\partial^2 \phi}{\partial t^2} - \vec{\nabla}^2 \phi + m^2 \phi &= 0 \rightarrow (1) \\ (1) \times (-i\phi^*) &\Rightarrow -i\phi^* \frac{\partial^2 \phi}{\partial t^2} + i\phi^* \vec{\nabla}^2 \phi + m^2 (-i\phi^* \phi) = 0 \rightarrow (2) \\ (1)^* \times (-i\phi) &\Rightarrow -i\phi \frac{\partial^2 \phi^*}{\partial t^2} + i\phi \vec{\nabla}^2 \phi^* + m^2 (-i\phi \phi^*) = 0 \rightarrow (3) \\ (2) - (3) &\Rightarrow \left( -i\phi^* \frac{\partial^2 \phi}{\partial t^2} + i\phi \frac{\partial^2 \phi^*}{\partial t^2} \right) \\ &\quad + (i\phi^* \vec{\nabla}^2 \phi - i\phi \vec{\nabla}^2 \phi^*) = 0 \rightarrow (4) \end{aligned}$$

Let us discuss this further we will consider  $\square$  square write it in many components  $\square$  square by  $\square t$  square. So, now, i work in units of  $c$  equal to one square  $\phi$  minus grad square  $\phi$  plus  $m^2$  square  $\phi$  equal to 0, change the wave function notation from  $\phi$   $\psi$  to  $\phi$  I mean you know you know particular need to do that, but let me denoted by  $\phi$ , because that is usual the standard way of denoting the fields who satisfy the Klein Gordon equation for the wave functions who satisfy the Klein Gordon equation.

Now let me multiply this by minus  $i\phi^*$ . So, let me called this equation 1. So, 1 times minus  $i\phi^*$  will give me minus  $i\phi^*$ , second derivative with respect to time of  $\phi$  minus  $i$ . So, plus  $i\phi^*$  grad square  $\phi$  minus  $i\phi^*$   $m^2$  in the beginning  $m$  outside.

So,  $m^2 \psi - i \psi^* \dot{\psi} = 0$ , let me call this equation 2. So, let me take the complex conjugate of this and multiply this by  $-i$ . So, that will give me  $-i \psi^* \ddot{\psi} = m^2 \psi^* \dot{\psi}$ . I have taken the complex conjugate of equation 1 minus again that becomes  $+i \dot{\psi}^* \ddot{\psi} - m^2 \psi^* \dot{\psi} = 0$ . So, this is equation 3. Now this equation 2 minus 3 will give me all right let me write it here 2 minus 3  $m^2$  terms are identical. So, they will cancel out.

Now, there is a  $-i \psi^* \ddot{\psi}$  plus  $i \dot{\psi}^* \ddot{\psi}$  over  $\dot{\psi}^2$  ok. Now plus  $i \dot{\psi}^* \ddot{\psi}$  minus  $i \dot{\psi}^* \ddot{\psi}$  equal to 0.

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The image shows a handwritten derivation on a light background. It starts with the product rule for the time derivative of  $\psi^* \dot{\psi}$ :

$$\frac{\partial}{\partial t} (\psi^* \dot{\psi}) = \psi^* \ddot{\psi} + \dot{\psi}^* \dot{\psi}$$

Then, it isolates the  $\psi^* \ddot{\psi}$  term:

$$\psi^* \ddot{\psi} = \frac{\partial}{\partial t} (\psi^* \dot{\psi}) - \dot{\psi}^* \dot{\psi}$$

Below this, it says "first term in" and then shows equation (4) being substituted into equation (5):

$$(4) \Rightarrow -i \left\{ \frac{\partial}{\partial t} (\psi^* \dot{\psi}) - \dot{\psi}^* \dot{\psi} - \frac{\partial}{\partial t} (\dot{\psi}^* \psi) + \dot{\psi}^* \dot{\psi} \right\} =$$

$$-i \frac{\partial}{\partial t} \left\{ \psi^* \dot{\psi} - \dot{\psi}^* \psi \right\} \rightarrow (5)$$

Now if I take time derivative acting on  $\psi^* \dot{\psi}$  by  $\dot{\psi}$ , that will give me  $\psi^* \ddot{\psi}$  plus  $\dot{\psi}^* \dot{\psi}$  minus  $i \dot{\psi}^* \ddot{\psi}$  plus  $i \dot{\psi}^* \ddot{\psi}$  over  $\dot{\psi}^2$  there is only my sorry plus  $\dot{\psi}^* \dot{\psi}$  by  $\dot{\psi}$ . Similarly for the other term time  $\dot{\psi}^* \psi$  by  $\dot{\psi}$  also we can write in the same fashion.

We use this to rewrite equation number 4 as above. So, not that here we have a term  $\psi^* \ddot{\psi}$  plus  $\dot{\psi}^* \dot{\psi}$  minus  $i \dot{\psi}^* \ddot{\psi}$  plus  $i \dot{\psi}^* \ddot{\psi}$  over  $\dot{\psi}^2$  is the first term on the right hand side here, but then that is equal to total time derivative or right derivative acting on both  $\psi^*$  and  $\dot{\psi}$  minus.

So,  $\phi^* \square \phi$  is what we have in this equation here, the first term on the left hand side of equation 4 apart from the  $-i$  that is equal to time derivative acting on  $\phi^* \phi$  together minus this term extra time on the right hand side of the first equation here, which is  $\phi^* \dot{\phi}$ .

So, that when we put back in equation 4 here will give us  $-i \phi^* \square \phi$  by  $\dot{\phi}$  is replaced by  $\dot{\phi}$  acting on  $\phi^* \phi$ , minus  $\phi^* \dot{\phi}$  plus we have taken a  $-i$  from the first wound therefore, it will be a  $+i \phi^* \dot{\phi}$  will become a  $-i$  know. So, there is a minus sign there because  $-i$  is taken out and we have a gain this is  $\phi^* \dot{\phi}$  minus  $\phi^* \dot{\phi}$  plus  $\phi^* \dot{\phi}$ .

So, this is the first term in equation 4. So, this first term in equation 4 is equal to. So, let me only focus on the first term that will give me  $-i$  these 2 cans the second term on the fourth term cancel.

So, that will give me  $-i$  total derivative acting on  $\phi^* \phi$  minus  $\phi^* \dot{\phi}$ . Similarly so, let me call this equation 5 ok. So, now, go back to equation 4 the second time is  $i \phi^* \square \phi$  minus  $i \phi^* \square \phi$ .

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Similarly,

$$\begin{aligned} \vec{\nabla} \cdot (-i \phi^* \vec{\nabla} \phi) &= -i \vec{\nabla} \phi^* \cdot \vec{\nabla} \phi - i \phi^* \vec{\nabla}^2 \phi \\ + i [\phi \nabla^2 \phi^* - i \phi^* \vec{\nabla}^2 \phi] &= \left\{ \vec{\nabla} \cdot (-i \phi^* \vec{\nabla} \phi) - i \vec{\nabla} \phi^* \cdot \vec{\nabla} \phi \right. \\ &\quad \left. + \vec{\nabla} \cdot (i \phi \vec{\nabla} \phi^*) + i \vec{\nabla} \phi \cdot \vec{\nabla} \phi^* \right\} \\ &= -i \vec{\nabla} \cdot \left\{ \phi^* \vec{\nabla} \phi - \phi \vec{\nabla} \phi^* \right\} \end{aligned}$$

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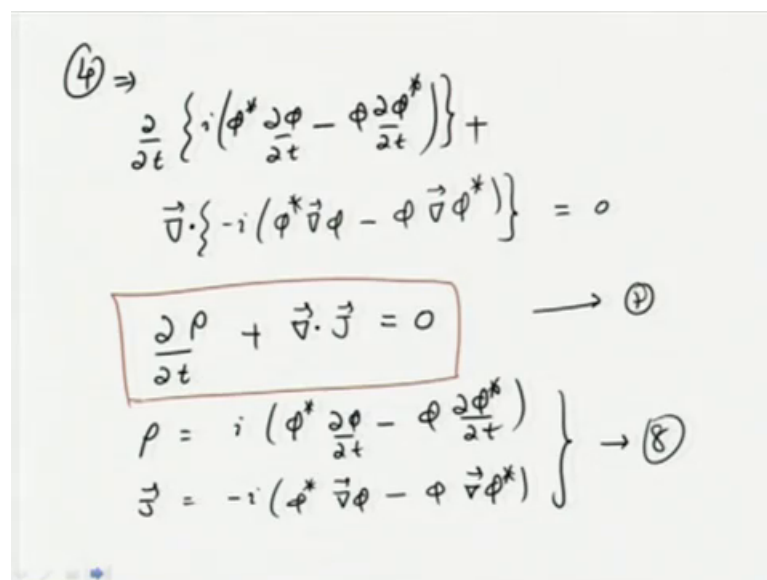


And in exactly similar way as we did in the case of the time derivative we can write grad gradient off minus i phi star grad phi dot here is equal to minus i grad phi star dot grad phi minus i phi star grad square phi.

So, the second term in equation 4 which is minus, which is plus I phi grad square phi star minus i phi star grad square phi can be written as del dot minus i phi star grad phi ok, minus minus i grad phi star dot grad phi. This is for the second time here a minus i phi star dot square phi and similarly for the other one, we have gain del dot i phi del phi minus minus of minus plus i phi sorry i del phi dot del phi star again the second and the fourth terms cancel.

And then we have this lhs or the rhs here is equal to. So, this is essentially equal to del dot minus i del dot phi star del phi minus phi del phi star ok. So, essentially what we have done is we have taken this equation 4 and written each of these term as time derivative for the first term acting on both phi star and dou phi by dou t time derivative is acting on both the terms both the factors and all there. And similarly for the gradient we have we written that in a more in a different fashion.

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$$\begin{aligned}
 (4) \Rightarrow & \frac{\partial}{\partial t} \left\{ i \left( \phi^* \frac{\partial \phi}{\partial t} - \phi \frac{\partial \phi^*}{\partial t} \right) \right\} + \\
 & \vec{\nabla} \cdot \left\{ -i \left( \phi^* \vec{\nabla} \phi - \phi \vec{\nabla} \phi^* \right) \right\} = 0 \\
 & \boxed{\frac{\partial \rho}{\partial t} + \vec{\nabla} \cdot \vec{J} = 0} \quad \rightarrow (7) \\
 & \left. \begin{aligned} \rho &= i \left( \phi^* \frac{\partial \phi}{\partial t} - \phi \frac{\partial \phi^*}{\partial t} \right) \\ \vec{J} &= -i \left( \phi^* \vec{\nabla} \phi - \phi \vec{\nabla} \phi^* \right) \end{aligned} \right\} \rightarrow (8)
 \end{aligned}$$

Why we are doing it because when we write it so, we have essentially equation 4 will give you time and derivative i star phi star dou phi by dou t minus phi phi star over dou t ok.

So, this is basically what we have from equation 5 here and similarly from let me call this equation 6 from 6 we have plus del dot minus i phi star grad phi minus phi grad phi star.

So, equation 4 can be written as time derivative of some term plus divergence of something else is equal to 0 I will write this in a compact way as dou by dou t of some rho plus divergence of some J equal to 0. So, this is 1 thing and now where rho is essentially as is rho is equal to i phi star dou phi by dou t minus phi dou phi star over dou t and j equal to minus i phi star grad phi minus phi grad phi star.

So, this thing here is a very compact notation compact equation you must have seen such equation in the earlier discussions at different under different topics, this is what is called the continuity equation ok.

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$$j^\mu = (\rho, \vec{j})$$

$$\frac{\partial \rho}{\partial t} + \vec{\nabla} \cdot \vec{j} = 0$$

$$\boxed{\partial_\mu j^\mu = 0}$$

Continuity eqn:

$$x^\mu x_\mu = x^0 x_0 - \vec{x} \cdot \vec{x}$$

$$\partial^\mu \partial_\mu = \partial_0^2 - \vec{\nabla}^2$$

$$\partial_\mu = \left( \frac{\partial}{\partial t}, +\vec{\nabla} \right)$$

$$j^\mu = (\rho, \vec{j})$$

This before we come to that interpretation of this i can actually write more compact current j mu which component rho and j vector and the continuity equation dou by dou t of rho plus divergence of j then becomes equal to 0, then becomes dou mu j mu equal to 0.

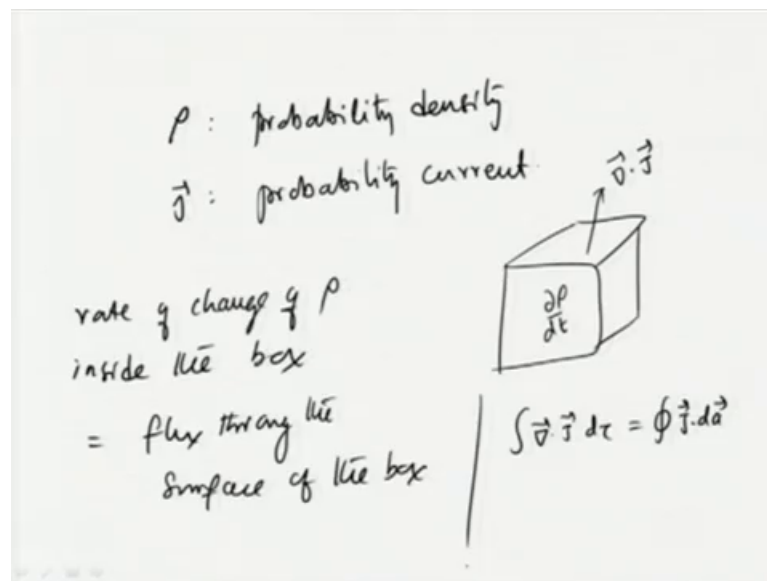
He notice that between the special and the temporal components here the 0th component and a 1 2 3 components the sign is not minus, but plus just as an aside. The x mu x mu

we had  $\partial_\mu \psi = 0$  minus  $\psi \cdot \partial_\mu \psi$  this was also the case, where we had the operator  $\partial_\mu$  equal to  $\partial_0$  square minus grad square relative sign.

But here we do not have that relative sign; because  $\partial_\mu$  is  $\partial/\partial t$  plus grads the covariant vector has a plus sign in this case. And  $j_\mu$  is a contravariant vector which also has a plus sign the way we have defined  $j$ .

So, this is the continuity equation written in a compact way the physical interpretation is that all right. So, this is the continuity equation. If you look at the expression here in equation 8,  $j$  is something which we have already familiar with in the case of non-relativistic quantum mechanics ok. So, there  $j$  is basically considered as the probability current.

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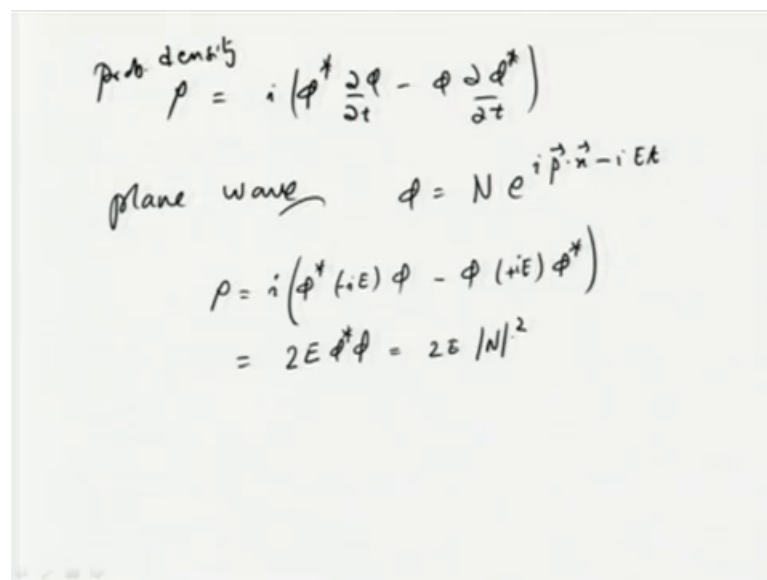
So, this here again we can interpret  $\rho$  as the probability density or the rate of  $\rho$  is the probability density and  $j$  is the probability current. The interpretation is this that if you consider a small volume say a small cube rate of change of probability density inside the volume is the same as the rate of change or the divergence of the  $J$ , which essentially is because of the divergence theorem or the Gauss's theorem this corresponds to the flux going out of the surface from the volume you see. So, minus of the scene is actually the flux which is going out of this volume.

So, it essentially changes the says inverse that rate of change of the continuity equation will tell you the rate of change of probability density rho inside the box is equal to flux flux is j dot the area.

So, you have to use the gauss's theorem divergence of a vector over integrate integrated over a volume is the same as J dot da a surface integral of the same vector over the closed surface which encloses the volume over the boundary of the volume. So, that is basically the flux the right hand side here is the flux of this thing of the field.

So, here it is the flux to the surface of the box and what flux is this the flux is basically the probability flux. So, it is it says for example, if you consider some number of particles inside this one, then it will actually tell you the rate of change of that minimum number of particles is equal to the number of particles that flows out or the probability for the number of particles 2 floors a kind of a statistical statement that is what we expect in quantum mechanics, but it essentially is very similar to the continuity equation that we have in any other physical situation like fluid dynamics water flow liquid flow all right.

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Handwritten mathematical derivations:

$$\text{prob density } \rho = i \left( \phi^* \frac{\partial \phi}{\partial t} - \phi \frac{\partial \phi^*}{\partial t} \right)$$

$$\text{plane wave } \phi = N e^{i \vec{p} \cdot \vec{x} - i E t}$$

$$\rho = i \left( \phi^* (-i E) \phi - \phi (i E) \phi^* \right)$$

$$= 2 E \phi^* \phi = 2 E |N|^2$$

Now, so we have the expression for rho equal to i phi star time derivative of phi minus phi time derivative of phi star this is the probability density. So, the probability density, if i take a plane wave solution or plane wave for the phi, then I can write it as e pi e power i P dot x minus i energy times time this is a plane wave.

Then rho is nothing but i phi star remains as 5 star dou phi by dou t will take out minus i and E, then phi itself minus phi as it is phi star now will have n star exponential minus i P dot x plus i E t. So, when I take the time derivative it is plus ie that comes out and then phi star as it is this is.

So, there is an I here overall factor and then there is an minus i E in the first term together that will give me a plus E. So, plus e and the second term again I have a minus sign. So, minus i E inside the bracket and an i outside. So, that is again going to give you a plus E. So, total 2 E phi star phi, which is 2 E en square mod N square.

So, rho in the case of plane wave solution is twice the energy of the solution wave function and or energy of the particular corresponding to the wave function, and the normalization a square all right.

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Handwritten text on a slide:

$$\phi = N e^{i\vec{p}\cdot\vec{x} - iEt} \text{ is a solution of}$$

$$\frac{\partial^2 \phi}{\partial t^2} - \nabla^2 \phi + m^2 \phi = 0$$

then  $\phi' = N e^{i\vec{p}\cdot\vec{x} + iEt}$  is also a solution

$$E^2 = \vec{p}^2$$

So, now, let us consider the 5 izzard is a phi equal to N E power i P dot x i P dot x minus i E t is a solution of Klein Gordon equation dou square phi by dou t square minus del square phi plus m square phi equal to 0, if this is all then phi equal to some other phi or phi prime.

So, let me take a yeah phi prime N e power i P dot x plus i E t this is also a solution. That is because the equation Klein Gordon equation is quadratic equation it actually second order derivatives are coming in.

So, for  $i$  plus  $i$  square and minus  $i$  square both will give you the same minus  $E$  square. So, essentially this is coming down to the situation where, we have to start with we have  $p$  square minus let me write on both of this actually give you, when you put it back in kg equation Klein Gordon equation both of this will give you,  $E$  square  $\phi$  minus  $P$  square  $\phi$  sorry there is already a minus sign there plus.

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Handwritten mathematical derivation showing solutions to the Klein-Gordon equation:

$$\phi = N e^{i\vec{p}\cdot\vec{x} - iEt} \text{ is a solution of } \frac{\partial^2 \phi}{\partial t^2} - \nabla^2 \phi + m^2 \phi = 0$$

$$\text{Then } \phi' = N e^{i\vec{p}\cdot\vec{x} + iEt} \text{ is also a solution}$$

Both will give,

$$-E^2 \phi + \vec{p}^2 \phi + m^2 \phi = 0$$

$$E^2 = \vec{p}^2 + m^2$$

$$E = \pm \sqrt{\vec{p}^2 + m^2}$$

So, this is minus this is plus this is plus  $m$  square  $\phi$  equal to 0 both of this will give you this.

So, this is essentially  $E$  square is equal to  $P$  square plus  $m$  square, which is basically the energy momentum relation relativistic energy momentum relation. And if I write down the expression for  $E$  I can actually write it as  $P$  square plus  $m$  square under root plus or minus.

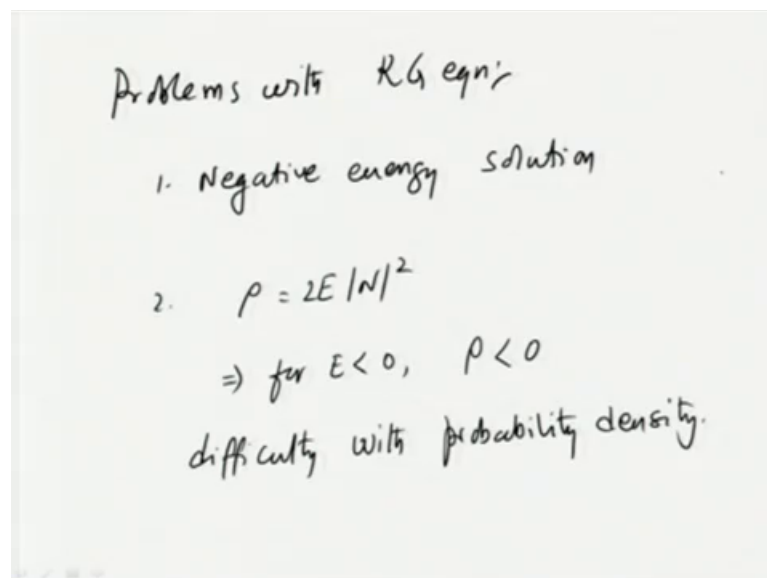
Now you can see that more clearly if whether I take plus  $E$  or minus  $E$  the equation should give the same result or the other way of saying it is that,  $E$  is a solution to this thing minus  $E$  is also a solution I mean expression with plus  $e$  is a solution to the kg equation expression with minus  $E$  is also a solution to the kg equation.

When we interpret  $E$  as energy of the particle corresponding to the wave function  $\psi$  this then put us in a difficult situation, it tells you that you have a solution with energy  $E$  for this particle.

And you can order in that case-avoid a solution another solution with minus E. So, a negative energy solution emerges comes along with the positive energy solution. This is a problem, because we are talking still about free particles they are not put any potential here these are all free particle solutions.

We do not have any extra potential that is why E square is equal to P square plus m square only the kinetic energy term and the mass term are there in the energy. So, negative energy is not a physically viable case we do not want to have negative energy solutions, but the Klein Gordon equation actually give you solutions with negative energy. So, this is one problem.

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So, problems with kg equation as it stands now are 1 negative energy solution. 2 for that let us look at the probability density in our case the probability density is equal to twice the energy times N square N square is not positive negative 2 is not negative. So, if E is negative then rho is negative.

So, that will give you for E less than 0 rho less than 0 how do we interpret rho then as probability density. So, difficulty with interpreting or let me simply write it as probability density you cannot have negative probabilities which is physically not valid.

So, these 2 problems are to be addressed to understand the relativistic equation a relativistic dynamics of particles through Klein Gordon equation we will come to how we actually address these questions in the next lecture.