

Statistical Thermodynamics for Engineers
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Lecture 07
Bose Einstein and Fermi Dirac Statistics

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9:41 AM Tue 9 Jan Untitled Notebook (31)

$W > W_{mp} > \bar{W}$
↓
↘

Total # of
microstates
mean # of
microstates

$N = 10^{23}, M = 2$

$W = M^N = 2^N \quad \bar{W} = \frac{2^N}{N+1}$

So, welcome to lecture 7 of the Statistical Thermodynamics for Engineers. So, just a brief recap. If you recall what we did in the last lecture is that W is greater than the most probable that greater than \bar{W} , W was the total number of microstates, these are the total number of microstates of a system and this is the mean of microstate. So, this was already there now, if your N is 10 to the power 23 and M is equal to 2. So, W that is the total number of microstates is given as M to the N then 2 the power N , then \bar{W} , the mean number of microstates is basically 2 to the power of N divided by N plus 1.

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$N = 10^{23}, M = 2$
 $W = m^N = 2^N \quad \bar{w} = \frac{2^N}{N+1}$
 $\ln W = N \ln 2 \approx 7.0 \times 10^{22}$
 $\ln \bar{w} = N \ln 2 - \ln(N+1) \approx 7.0 \times 10^{22} - 53$
 $\ln \bar{w} \approx \ln W$
 $\ln W \approx \ln W_{mp} \approx \ln \bar{w}$

Now, if we now look at it carefully, $\ln W$ that is if we take the natural log of W , which will be in $N \ln 2$, which roughly translates to about 7.0 into 10 to the power 22, $\ln \bar{w}$ which is basically the logarithm mean microstates and up to minus plus 1, which also translates to 7.0 into 10 to the power 22 minus 53. So, what we can see here is $\ln \bar{w}$ is almost the same as in W , because you can see that this particular number is so large compared to this, that these two are basically one in the same. So, in other words, this also means that $\ln W$ is therefore almost equal to \ln most probable, which is almost equal to $\ln \bar{w}$.

So, all it means is that the logarithmic of the total number of microstates is almost the same as the most probable number, the number of microstates associated with the most probable microstate and that is almost equal to the mean number of microstates. So, that is eliminating.

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$\lim_{N \rightarrow \infty} \frac{\ln W_{mp}}{\ln W} = 1$
 It seems that for macroscopic thermodynamic systems almost all microstates are associated with the most probable macrostate.

So, what we can say is that in the limit as N approaches infinity, \ln of W_{mp} divided by $\ln W$ is almost equal to one or in other words, it seems that for macroscopic thermodynamics that was to be thermodynamic systems comma almost all microstates are associated with the most probable microstate. So, this is the most important inference that because this is true, and because N is pretty large, it seems that for macroscopic thermodynamic systems, almost all microstates. This is the most important part almost all microstates are associated with the most probable microstate, all microstates are associated with the most probable microstate. So, the most probable distribution must represent the equilibrium particle distribution.

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It seems that
Thermodynamic systems
almost all microstates
are associated with the
most probable macrostate
↓
represents the equilibrium
particle distribution

So, this therefore, represents the equilibrium particle distribution. So, this therefore represents the equilibrium particle distribution. So, that is eliminated to begin with.

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Bose-Einstein and
Fermi-Dirac Statistics

BE statistics

↳ In distinguishable particles
with no limit on the
of particles per energy
state

So, let us now look at the first level of statistics because we have now we have covered the different aspects of macrostate microstate we know a lot distinguishable indistinguishable particles also. So, we are going to look at Bose Einstein, Einstein and Fermi Dirac statistics. So, first let us look at the Bose Einstein statistics or the BE statistics. So, the BE statistics that is the Bose Einstein statistics after S.N. Bose and Albert Einstein. So, this the premise is that indistinguishable particles with no limit on the number of particles per energy state.

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In distinguishable particles with no limit on the # of particles per energy state

Let us consider 1 energy level

of ways in which N_j bosons in a single energy level E_j maybe distributed in g_j energy states

degeneracy

So, let us consider do it slowly one energy level and let us consider the number of ways in which N_j bosons which is the number of particles a single energy level E_j maybe distributed g_j . So, we consider 1 energy level and we are going to see the number of ways in which N_j bosons because both sides test statistics that bosons so, these are the number of results in a single energy level E_j may be distributed in g_j energy states where g_j is the degeneracy. So, you have a physically g_j compartments you have N_j indistinguishable particles there is no limit on the number of particles or energy state. So, 1 energy state can have as many number of particles as you want.

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N_j identical, indistinguishable objects can be arranged in g_j containers with no limit in # of objects per container

degeneracy

So, we have to consider that how this distribution is possible. So, this is the problem is therefore, equivalent to N_j identical indistinguishable objects we already have done this objects can be arranged in g_j containers it is like what we knew as M containers with with no limit in the number of objects per container.

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$W_j = \frac{(N_j + g_j - 1)!}{N_j! (g_j - 1)!}$

Now since each energy level is independent total # of ways of obtaining an arbitrary particle dist'n

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$W_j = \frac{(N_j + g_j - 1)!}{N_j! (g_j - 1)!}$

Now since each energy level is independent total # of ways of obtaining an arbitrary particle dist'n

$W_{BE} = \prod_j W_j$
 $= \prod_j \frac{(N_j + g_j - 1)!}{N_j! (g_j - 1)!}$

So, therefore, we write it as W_j is equal to N_j plus g_j minus 1 factorial divided by N_j factorial g_j minus 1. So, these are the number of arrangements that we have. So, since each energy level is independent, so, now, since, each energy level is independent because it does not depend on other one, each energy level is independent. This is important, independent. The total number of ways obtaining an arbitrary particle distribution, write it as W_{BE} as Bose Einstein is basically will be the multiplicative of all the j energy levels that W_j which

essentially translates to j, N_j plus g_j minus 1 factorial divided by N_j factorial g_j minus 1 factorial.

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$$W_{BE} = \prod_j \frac{(N_j + g_j - 1)!}{N_j! (g_j - 1)!}$$

W_{BE} : generic # of microstates per macrostate for BE statistics

So, this is this is the, this is essentially what the distribution should look like. So, WBE is basically the generic number of microstates per macrostate for BE statistics, the generic number of microstates per macrostate for Bose Einstein statistics, they got the point that these are indistinguishable particles, they are identical and how they can be distributed in g_j containers, there is n_j number of particles in energy levels. So, what do you do after this?

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W_{BE} : generic # of microstates per macrostate for BE statistics

$$\ln W_{BE} = \sum_j \{ \ln(N_j + g_j)! - \ln N_j! - \ln g_j! \}$$

$g_j \gg 1 \rightarrow$ degeneracy is usually a large #

We follow the protocol that means, we take the logarithm of WBE, which therefore, converts this multiplicative sign into an additive sign. So, this becomes $\ln N_j$ plus g_j factorial. So, these are all the brackets by the way so, and then N_j factorial minus $\ln g_j$. So, g_j which was greater than 1 because we said that degeneracies usually a large number, degeneracy is usually large number.

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$$j - \ln N_j!$$

$$- \ln g_j!$$

$$g_j \gg 1 \rightarrow \text{degeneracy is usually a large \#}$$
 use Stirling's approx

$$\ln N! = N \ln N - N$$

$$\ln W_{BE} = \sum_j \{ (N_j + g_j) \ln(N_j + g_j) - N_j \ln N_j - g_j \ln g_j \}$$

So, use the Stirling's approximation which is basically $\ln N$ factorial is equal to $N \ln N$ minus N , so this is Stirling's approximation when you apply it to this particular context $\ln W_{BE}$ summation j , N_j plus g_j , $\ln N_j$ plus g_j minus N_j , $\ln N_j$ minus $g_j \ln g_j$.

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$$\ln W_{BE} = \sum_j \{ (N_j + g_j) \ln(N_j + g_j) - N_j \ln N_j - g_j \ln g_j \}$$

$$\ln W_{BE} = \sum_j \left\{ N_j \ln \frac{g_j + N_j}{N_j} + g_j \ln \frac{g_j + N_j}{g_j} \right\}$$

Now, simplifying it a little bit further so, this gives you $j N_j g_j$ plus $\ln g_j$ plus N_j by N_j plus $g_j \ln g_j$ plus N_j by g_j . So, this will be your Bose Einstein statistics. So, this up to this particular point worked out because this is what the statistics would look like but we are not done yet.

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$$\ln W_{BE} = \sum_j \left\{ \ln \frac{N_j!}{g_j^{N_j}} + g_j \ln \frac{g_j + N_j}{g_j} \right\}$$

Fermi-Dirac Statistics
Fermions

N_j fermions in E_j energy level in g_j container with limit of 1 particle per container

So, we will do the Fermi Dirac, now take it up to the same level and then we will see most probable distributions as the Fermi Dirac statistics. Next is the Fermi Dirac statistics. So, these are valid for fermions, small fermions, so class of Fermions. So, again the same thing N_j fermions in E_j energy level in g_j containers for this time we have a limit with the limit of one particle per container.

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fermions

N_j fermions in E_j energy level in g_j container with limit of 1 particle per container

→ Identical and indistinguishable

$$W_j = \frac{g_j!}{N_j! (g_j - N_j)!}$$

↳ indistinguishability accounts for it

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$$W_j = \frac{g_j!}{N_j!(g_j - N_j)!}$$
 ↳ indistinguishability accounts for it

$$W_{FD} = \prod_j W_j = \prod_j \frac{g_j!}{N_j!(g_j - N_j)!}$$
 ↳ energy levels are independent generic # of microstates per macrostate

And they are, of course identical and indistinguishable so, W_j is equal to g_j factorial by N_j factorial into g_j minus N_j factorial obviously already know, this accounts for it, for the indistinguishability accounts for it, this already know as well and so, the Fermi Dirac statistics, therefore (16:20) down to the multiplicative of W_j 's which essentially translates to this g_j factorial divided by N_j factorial g_j minus N_j .

So, this is once again the generic number of microstates per macrostates also the same assumption valid because the energy levels this essentially translates to the energy levels are independent.

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↳ energy levels are independent generic # of microstates per macrostate

$$\ln W_{FD} = \sum_j \left\{ \ln g_j! - \ln N_j! - \ln (g_j - N_j)! \right\}$$
 Stirling's approx

$$\ln W_{FD} = \sum_j \left\{ g_j \ln \frac{g_j - N_j}{N_j} - g_j \ln \frac{g_j - N_j}{g_j} \right\}$$

So, now if you cast $\ln W_{FD}$, once again we take the log g_j factorial minus $\ln N_j$ factorial minus $\ln g_j$ minus N_j factorial, so this is the bracket. Again, use Stirling's approximation

which essentially translates $\ln W$ is equal to summation $j N_j \ln g_j$ minus N_j divided by minus N_j minus $g_j \ln g_j$ minus N_j by g_j . So, this will be Fermi Dirac statistics. So, you saw how we use the statistical tools that we developed already generic situations and how it is applicable when you actually do the all side and the Fermi Dirac statistics.

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most probable distribution

$$\ln W_{\text{FD}} = \sum_j \left[N_j \ln \frac{g_j + N_j}{N_j} \mp g_j \ln \frac{g_j \pm N_j}{g_j} \right]$$

most probable distribution

So, now, we choose the most probable distribution that is what you want to find out. So, this $\ln W$, whether it is Bose Einstein or Fermi Dirac, it is like the distributions are like this is $N_j \ln$ depends on which distribution you are looking at this plus minus $g_j \ln g_j$ plus minus N_j by g_j , so, this is the distribution that you are looking at the most probable distribution in order to find that out, our task is to find out the most probable distribution. So, these are the Fermi Dirac statistics for any arbitrary distribution.

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$$\ln W_{\{ \omega E \}} = \sum_j \left\{ N_j \ln \frac{\omega_j}{N_j} \mp g_j \ln \frac{g_j \pm N_j}{g_j} \right\}$$

most probable distribution

Two constraints

$$\sum_j N_j = N \quad \text{--- mass}$$

$$\sum_j N_j \epsilon_j = E \quad \text{--- energy}$$

So, for that we need we introduced two constraints. The first one is that summation of N_j is equal to N , which is basically mass observation then you have sum over $N_j \epsilon_j$ is equal to E is basically the energy conservation. So, these are both constraints.

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Two constraints

$$\sum_j N_j = N \quad \text{--- mass}$$

$$\sum_j N_j \epsilon_j = E \quad \text{--- energy}$$

$$\ln W = \sum_j \left\{ N_j \ln \left(\frac{g_j \pm N_j}{N_j} \right) \mp g_j \ln \frac{g_j \pm N_j}{g_j} \right\}$$

$$d \ln W = \sum_j \left\{ \ln \left(\frac{g_j \pm N_j}{g_j \pm N_j} \right) \pm \frac{N_j}{g_j \pm N_j} - \ln N_j - 1 + \frac{g_j}{g_j \pm N_j} \right\} dN_j$$

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$$d \ln W = \sum_j \left\{ \ln(g_j \pm N_j) \pm \frac{N_j}{g_j \pm N_j} - \ln N_j - 1 + \frac{E_j}{g_j \pm N_j} \right\} dN_j$$

$$d \ln W = \sum_j \left\{ \ln(g_j \pm N_j) - \ln N_j \right\} dN_j$$

g_j and E_j are taken as constants during differentiation

$g_j \rightarrow$ integer

$E_j \Rightarrow f(\text{total vol. of the system}) \rightarrow$ constant

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$g_j \rightarrow$ integer

$E_j \Rightarrow f(\text{total vol. of the system}) \rightarrow$ constant for an isolated system

$$\therefore \sum_j dN_j = 0$$

$$\sum_j E_j dN_j = 0$$

So, we saw that now if we once again $\ln W$ is summation $N_j \ln g_j$ minus plus minus N_j minus $N_j \ln N_j$ plus minus $g_j \ln g_j$ minus plus $g_j \ln g_j$, so this is the expression. What you do is that you differentiate us because you have to find out the most probable distribution. So, once you do that you get $\ln g_j$ plus minus $N_j dN_j$ so, that is what we have done so, if you now do it further $\ln g_j$ plus minus, you work out the steps, so these are the things that you get. Now, g_j and E_j are constants.

So, we have taken g_j and E_j are taken as constants during the differentiation. So, g_j is an integer this degeneracy and E_j is of course it is basically a function of the total volume of the system and it is constant for an isolated system. Therefore, summation dN_j is equal to 0, summation $E_j dN_j$ equal to 0.

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system) \rightarrow constant for an isolated system

$$\therefore \sum_j dN_j = 0$$

$$\sum_j \epsilon_j dN_j = 0$$

Introduce Lagrange multipliers

$$\sum_j \{ \ln(g_j \pm N_j) = \ln N_j - \alpha - \beta \epsilon_j \} dN_j = 0$$

α, β are Lagrange multipliers

So, now we have the situation and you know so, introducing the concept of Lagrange multipliers now, introduce Lagrange multipliers and what you have is summation $j \ln g_j$ plus minus N_j minus $\ln N_j$ minus α minus $\beta E_j dN_j$ equal to 0. Alpha and beta are Lagrange multipliers.

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Introduce

$$\sum_j \{ \ln(g_j \pm N_j) = \ln N_j - \alpha - \beta \epsilon_j \} dN_j = 0$$

α, β are Lagrange multipliers

This can be achieved for all j iff

$$\ln \frac{g_j \pm N_j}{N_j} = \alpha + \beta \epsilon_j$$

$$\therefore N_j = \frac{g_j}{\exp(\alpha + \beta \epsilon_j) \mp 1}$$

This particular thing can only be achieved for all j and only if $\ln g_j$ plus minus N_j divided by N_j is equal to α plus βE_j . So, therefore, N_j is equal to g_j divided by exponential α plus βE_j minus plus 1.

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$$\ln \frac{g_j \pm N_j}{N_j} = \alpha + \beta \epsilon_j$$

$$N_j = \frac{g_j}{\exp(\alpha + \beta \epsilon_j) \mp 1}$$

most probable
distribution and for
large N it is the
only distribution.

Entropy and Equilibrium
particle distribution

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$$\sum_j \epsilon_j dN_j = 0$$

Introduce Lagrange multipliers

$$\sum_j \{ \ln(g_j \pm N_j) - \ln N_j - \alpha - \beta \epsilon_j \} dN_j = 0$$

α, β are Lagrange multipliers

This can be achieved for all j if

$$\ln \frac{g_j \pm N_j}{N_j} = \alpha + \beta \epsilon_j$$

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$$d \ln \Omega = \sum_j \{ \ln(g_j \pm N_j) - \ln N_j \} dN_j$$

g_j and ϵ_j are taken as constants during differentiation

$g_j \rightarrow$ no. of states

$\epsilon_j \Rightarrow$ total vol. of the system \rightarrow constant for an isolated system

$$\sum_j d \ln \Omega = 0$$

$$\sum_j dN_j = 0$$

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This will be the most probable distribution and as we saw is only distribution and for large it is for large N , it is the only distribution. So, as you can see now, so, this is depending on which one you are looking at this is the Bose Einstein is the most probable distribution this is how N_j particles are distributed in g_j energy level, so, this is this is another interesting thing and just to recap our data Lagrange multipliers, this is valid as you can see why, E_g is basically the function of a system that is constant system and that comes from Formula macroscopic definition and we have done a lot of stuff using Stirling's approximation stuff like that to arrive at this particular session.

So, the next class we are going to look at the next lecture we are going to look at entropy and the equilibrium particle distribution. So, coming up is entropy and equilibrium distribution. So, this is what is coming up next, in the next lecture, which will be lecture 8.

And there we are going to look at how the most probable distribution can be found out, but here in this case you can see we possibly do much more alpha or beta is an unknown. So, therefore, this distribution that whatever I have given you is incomplete. This is incomplete. Since have been, so hence they are hence you cannot really have a fair idea of what is your N_j to be.

So, N_j if it is the most probable distribution, we still need to see what is this alpha and beta all about. Because only then you can find out the most probable distribution and for that we need entropy and we need a few other things, which we are going to cover in the next lecture. So, thank you.