

**Quantum Mechanics and Molecular Spectroscopy**  
**Prof. G. Naresh Patwari**  
**Department of Chemistry**  
**Indian Institute of Technology - Bombay**

**Lecture – 24**  
**Selection Rules**

Hello, welcome to the lecture number 24 of the course quantum mechanics and molecular Spectroscopy okay. Now in the several classes, we were looking at the transition moment integral and its relation to other quantities such as Einstein's A coefficient, B coefficient and the absorption spectrum through Beer-Lambert law, but from now onwards I am going to take a slightly different turn okay and try to get the selection rules okay.

**(Refer Slide Time: 01:00)**

$$P_f(t) = \frac{E_0^2}{4t^2} \left( \frac{\omega_{fi}}{\omega} \right) \frac{\sin^2 \left( \frac{\Delta\omega}{2} t \right)}{\left( \frac{\Delta\omega}{2} \right)^2} |\langle f | \mu \cdot \hat{\epsilon} | i \rangle|^2$$

↳ Transition moment integral.

$$P_f(t) \propto |\langle f | \mu \cdot \hat{\epsilon} | i \rangle|^2$$

in the vicinity of  $\omega_f$

$$P_f(t) = \frac{E_0^2}{4t^2} \frac{\sin^2 \left( \frac{\Delta\omega}{2} t \right)}{\left( \frac{\Delta\omega}{2} \right)^2} |\langle f | \mu \cdot \hat{\epsilon} | i \rangle|^2 \Rightarrow \text{Line shape.}$$

Now if you go back quite a few classes, then we came across this equation which I will write it down as probability of transition P of t to state f is given by E 0 square divided by 4 h bar square omega fi by omega square sin square delta omega by 2 into t divided by delta omega by 2 square modulus of f mu dot e i square and we remember we called it as transition moment integral okay.

Now it turns out that Pf of t is proportional to your transition moment integral square. So what are you doing here? You are going from a state i to state f okay and we also related all this to Einstein's A coefficient, Einstein B coefficient and the extinction coefficient through Beer-Lambert law. We also looked at the line shapes okay. All these experimental quantities could be related to this.

Now it turns out that if I slightly rewrite, so in the vicinity of  $\omega \approx \omega_{fi}$  okay that means what I am doing is  $\omega \approx \omega_{fi}$  is approximately equal to  $\omega_{fi}$  that means you will have a transient only when there is a resonance okay. So in that scenario, one can rewrite this equation Pf of t, so when you have  $\omega \approx \omega_{fi}$  is equal to  $\omega_{fi}$ , I can ignore this term, so I will get this will be equal to  $E_0^2 \sin^2(\frac{\Delta\omega}{2} t) / (\frac{\Delta\omega}{2})^2$  okay.

Now we know this function is the line shape okay, it determines the line shape okay. Now what we have? Of course, all these are non zeros, so the selection will be allowed or not is given by this integral okay.

**(Refer Slide Time: 04:16)**

So one can write transition moment integral  $TMI = \langle f | \mu \cdot \epsilon | i \rangle$  okay. This is my transition moment integral and whether the transition from 1 to 2 or f to i will be allowed or not will depend on this integral. If this integral is 0, the transition is forbidden. If this is 0 then it is forbidden, this is not equal to 0, then you are allowed and 0 not equal to 0 they combine what is known as selection rules okay.

Now what we will do in this lecture is to look at this operator  $\mu$  okay. Now let us think of a molecule AB okay. So two atoms A and B are combined and A and B because they are not the same, so there will be say let us say  $\delta^+$  and  $\delta^-$ , so you have a dipole and if the equilibrium distance is  $r_0$ , so for example if you have AB molecule, one can draw a potential energy surface like this that is the zero of energy.

The distance goes to infinity asymptotically okay. Now what I am trying to do is that, of course if you go from here to say here there is a huge change in the geometry or the bond distance that it spans. So one can always think that the movements are not, so one is looking at around the equilibrium position if you look at around the equilibrium geometry, one can write  $\mu$  as a Taylor series expansion okay.

So, what I will write  $\mu = \mu_0 + d\mu$  by  $dr$  evaluated  $r_0$  into  $r + d^2\mu$  by  $dr^2$  evaluated  $r_0$  into  $r^2 + d^3\mu$  by  $dr^3$  evaluate at  $r_0$  into  $r^3 + \text{etc.}$ , etc. okay, one can write a Taylor series. Now let us look at it in a slightly more useful way.

**(Refer Slide Time: 08:23)**

The image shows handwritten mathematical work. At the top, it shows the Taylor series expansion of the dipole moment  $\mu$  around the equilibrium geometry  $r_0$ :

$$\mu = \mu_0 + \left. \frac{\partial \mu}{\partial r} \right|_{r_0} \cdot r + \frac{1}{2!} \left. \frac{\partial^2 \mu}{\partial r^2} \right|_{r_0} \cdot r^2 + \frac{1}{3!} \left. \frac{\partial^3 \mu}{\partial r^3} \right|_{r_0} \cdot r^3 + \dots$$

An arrow points from  $\mu_0$  to the text "Permanent Dipole". Below this, the expansion is written in a simplified form:

$$\mu = \mu_0 + \alpha \cdot r + \beta r^2 + \gamma r^3 + \dots$$

The coefficients are defined as:

$$\alpha = \left. \frac{\partial \mu}{\partial r} \right|_{r_0} \quad \beta = \frac{1}{2} \left. \frac{\partial^2 \mu}{\partial r^2} \right|_{r_0} \quad \gamma = \frac{1}{6} \left. \frac{\partial^3 \mu}{\partial r^3} \right|_{r_0}$$

At the bottom, the expansion is enclosed in a box and labeled as the Taylor series around the equilibrium geometry:

$$\text{Taylor series around the equilibrium geometry} = \left( \mu_0 + \alpha \cdot r + \beta r^2 + \gamma r^3 + \dots \right)$$

So the dipole moment  $\mu$  is a Taylor series expansion some fixed  $+ d\mu$  by  $dr$  at  $r_0$  multiplied by  $r + 1$  over  $2$  factorial  $d^2\mu$  by  $dr^2$  evaluated  $r_0$  into  $r^2 + 1$  over  $3$  factorial  $d^3\mu$  by  $dr^3$  evaluated  $r_0$   $r^3 + \text{etc.}$  So this is how I can write my  $\mu$  as a Taylor series expansion around the equilibrium geometry and this  $\mu_0$  is the permanent dipole and what is  $d\mu$  by  $dr$ , how the dipole moment changes with respect to  $r$  okay.

This is the first derivative and you have second derivative. So one can write it as  $\mu = \mu_0 + \alpha \cdot r + 1$  over  $2$   $\beta r^2 + 1$  over  $6$   $\gamma r^3$ , etc., etc. where  $\alpha$  equals to  $d\mu$  by  $dr$  evaluated  $r_0$ ,  $\beta$  is nothing but  $d^2\mu$  by  $dr^2$  evaluated  $r_0$  and  $\gamma$  equals to  $d^3\mu$ , actually one could take even the constants into, so if I write this as half then I can delete this and if I delete this then one can write  $1$  over  $6$   $d^3\mu$  by  $dr^3$  okay.

So one can write such equation. So this is your permanent dipole, this is a first derivative of mu, this is second derivative coefficient and the third derivative coefficients okay. So, this is how one can write in terms of Taylor series expansion. So therefore, your TMI around the equilibrium geometry is given by  $f \mu_0 + \alpha \dot{r} + \beta r^2 + \gamma r^3 + \dots$ , etc. So this is the transition moment that we need to evaluate okay.

When I say dipole moment, its basically linear combination of permanent dipole moment and its derivative at various levels of various degrees with respect to r okay. So, this is the transient moment integral that I need to consider.

**(Refer Slide Time: 12:04)**

$$TMI = \langle f | \mu_0 + \alpha r + \beta r^2 + \gamma r^3 + \dots | i \rangle$$

$$H^0 | i \rangle = E_i | i \rangle \quad H^1 | f \rangle = E_f | f \rangle \quad \text{KE of electrons}$$

$$H^0_{\text{mole}} = \underbrace{\frac{-\hbar^2}{2} \sum \frac{\nabla_i^2}{m_i}}_{\text{K.E of Nuclei}} + \underbrace{\frac{\hbar^2}{2m_e} \sum \nabla_i^2}_{\text{KE of electrons}} + \underbrace{\frac{1}{4\pi\epsilon_0} \sum \sum \frac{z_i z_j e^2}{R_{ij}}}_{\text{PE of c-c}} + \underbrace{\frac{1}{4\pi\epsilon_0} \sum \sum \frac{z_i z_p e^2}{R_{ip}}}_{\text{PE of N-N}} + \underbrace{\frac{1}{4\pi\epsilon_0} \sum \sum \frac{z_i z_j e^2}{R_{ij}}}_{\text{PE of c-N}}$$

So therefore, one can write  $TMI = f \mu_0 + \alpha r + \beta r^2 + \gamma r^3 + \dots$  so that is the transition moment integral I am interested in okay. Now let us look at this equation phi okay, so what are my solutions. So what are my solutions I said  $H_0$  into  $i = E_i i$  and  $H_0$  into  $f = E_f f$  okay, what are my  $H_0$ ,  $H_0$  is the Hamiltonian for which I know the solutions. So,  $E$  and  $E_f$  are the eigenvalues of the functions eigenfunctions  $i$  and  $f$  okay.

Now what is  $H_0$  that is the thing that we have to figure out okay. In this course from up till now we have never bothered what is  $H_0$ ,  $H_0$  we just assume is some Hamiltonian for which we know the solutions and once we know the solution we can evaluate the transient moment integral and that transient moment integral can relate to various experimentally measurable quantities, but right now we have not even thought about what the  $H_0$  could be?

Now if you write the entire molecular Hamiltonian which I am going to write so H molecule okay, so H0 molecule is nothing but  $-\hbar^2 \sum_{\alpha} \nabla_{\alpha}^2$  where  $\alpha$  is its index minus  $\hbar^2 \sum_i \nabla_i^2$  that is the kinetic energy of the nuclei where  $\alpha$  is its index minus  $\hbar^2 \sum_i \nabla_i^2$  so  $i$  is the index on the electron so the kinetic energy of all the electrons minus attraction between the nucleus and all the electrons.

So, this will be  $\frac{1}{4\pi\epsilon_0} \sum_{\alpha} \sum_{\beta} \frac{e^2}{R_{\alpha\beta}}$  so there is a distance between the electron and the nucleus, then you have a nuclear-nuclear repulsion so that is equal to  $\frac{1}{4\pi\epsilon_0} \sum_{\alpha} \sum_{\beta} \frac{e^2}{R_{\alpha\beta}}$  that is the nuclear repulsion because  $\alpha$  and  $\beta$  are 2 nucleus.

And you do not want to count the repulsion between the same nucleus that is why  $\alpha$  is not equal to  $\beta$  and then you have to have only one, if nucleus 1 is repelling nucleus 2 so it is same as nucleus 2 repelling nucleus 1, so to make sure that you do not repeat the terms use this kind of notation plus similarly with electron  $\frac{1}{4\pi\epsilon_0} \sum_i \sum_{j > i} \frac{e^2}{r_{ij}}$ . So, we have five terms, so this is KE of nuclei, this is KE of electrons.

This is nothing but PE of electrons and nucleus that is attraction, this is what PE of nucleus to nucleus that is repulsion and this is nothing but your PE of e that is nothing but electron-electron repulsion.

**(Refer Slide Time: 17:03)**

$H^0 \Psi(r, R) = E_n \Psi(r, R)$   
 Born-Oppenheimer approximation.  
 $\Psi(r, R) = \chi_n(r, R)$   
 $H^0 = H_n + H_e$   
 $H_n \chi_n = E_n \chi_n$   
 $E = E_n + E_e$   
 $PE = E_n + E_e$

The diagram shows a potential energy curve with a minimum. A horizontal line represents the total energy level, which is above the minimum of the potential well. The vertical axis is labeled 'E' and the horizontal axis is labeled 'r'.

$H_0$  of molecule multiplied by  $\psi$  now which is a function of electrons and nucleus should be equal to  $E_n \psi$  which is function of electrons and nucleus okay. Now it turns out that this is not a possible solution, so what we do is what we use something called Born Oppenheimer approximation okay.

This approximation says that the total wave function  $\psi$  of electrons and nucleus can be separated as  $\chi$  of nucleus and  $\phi$  of electrons parametrically dependent on nuclei and the total Hamiltonian  $H_0$  can be written as  $H_{\text{nuclei}} + H_{\text{electrons}}$  okay. Now it turns out that when I look at  $H_{\text{nuclei}}$ , the  $H_{\text{nuclei}}$  even though is this  $H_{\text{nuclei}}$  the total energy is equal to  $H_{\text{nuclei}} \chi_{\text{nucleus}}$  okay will give you  $E_{\text{nucleus}} \chi_{\text{nucleus}}$ .

But the total energy  $E$  is nothing but  $E_n + E_{\text{electron}}$ , but you see the  $\phi$ , so now if you have  $H_{\text{electron}}$  acting on  $\phi$  electron and that is parametrically depends on nucleus. What does parametrically depend on nucleus means, for every different nuclear position the Hamiltonian of the electrons will be different give a different solution. So, that will be nothing but  $E_e + \phi$ , so this  $E_e$  also depends on nucleus because the wave function itself.

So if you take the total energy or the nuclear energy, nuclear energy is just not the repulsion but it also kind of has the electronic component into it and that is called potential energy okay. So when you have PE that is nothing but you have  $E_n \text{ nucleus} + E_{\text{electron}}$  okay. So what you do is you can make a potential surface okay that is  $E$  so that depends on how the electrons and the nuclei are arranged as a function of distance okay.

**(Refer Slide Time: 20:17)**

$$\Psi(\rho) = \chi_n \cdot \Phi(\rho; \nu)$$

$$T_{M\hat{I}} = \langle [\chi_n \Phi(\rho; \nu)]_{\alpha, \beta, \gamma} \left| \frac{\hat{\mu} \cdot \hat{\xi}}{r} \right| [\chi_n \Phi(\rho; \nu)]_{\alpha, \beta, \gamma} \rangle$$

$$T_{M\hat{I}} = \langle [\chi_n \Phi(\rho; \nu)]_{\alpha, \beta, \gamma} \left| \frac{\mu_x + i\mu_y + \mu_z \hat{\sigma}_z}{r} \right| [\chi_n \Phi(\rho; \nu)]_{\alpha, \beta, \gamma} \rangle$$

Now one of the issues that we have is the following. So when you have total wave function which is when the total wave function  $\psi$  of  $e, n$  is nothing but  $\chi$  of  $N$  into  $\phi$  of  $e$  and parametrically depend on  $N$  okay, now this is my total solution okay, and of course it will have its own quantum numbers. So, what I am looking at is the following okay. I am starting from a final function  $f$ , now this  $f$  function could be written as the following.

$\chi$  nucleus  $\phi$  electron nucleus, all of this having some quantum number let us say  $a, b, c$  or I will call it as  $a'$ ,  $b'$ ,  $c'$ , I will come to what these are okay and this multiplied by your  $\mu \cdot E$  multiplied by function  $i$  which I will call it as  $\chi N \phi e \cdot N$  having quantum numbers  $a, b, c$  okay. So, what I am trying to do here I am going from initial wave function which will have quantum number  $a, b, c$ .

And this quantum number could involve the nuclear quantum numbers or quantum numbers related to the nuclear motion and the quantum numbers related to the electronic motion and with some values  $a, b, c$  and I am going from there and changing my quantum numbers to  $a'$ ,  $b'$ ,  $c'$  okay. Now what I am trying to do is the following. I am going from one set of electronic and nuclear coordinates to a different set of electronic and nuclear coordinates.

And this transition is brought about by this operator. So, now we can think of it this way, so this operator is really complicated in some sense is that  $\chi N \phi e, N$  whole thing of  $a'$ ,  $b'$ ,  $c'$   $\mu_0 + r + \beta r^2 + \gamma r^3$  whole  $\chi N$  to  $\phi e, N$  whole thing of  $a, b, c$ . So, I am looking at such a transition moment integral okay, so when I take the total wave function okay.

So this is how we are going to look at this transition moment integral, but of course the way it is written it is almost impossible to solve, so we have to make right approximations to be able to solve this transition moment integral, evaluate the transition moment integral and look at the selection rules which we will continue in the next lecture. Thank you.